

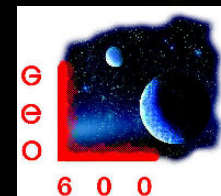
# Method for a safe statistical veto using IFO channels

LIGO-T070068-00-Z



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LSC/Virgo meeting, March 2007, Det-Char-session

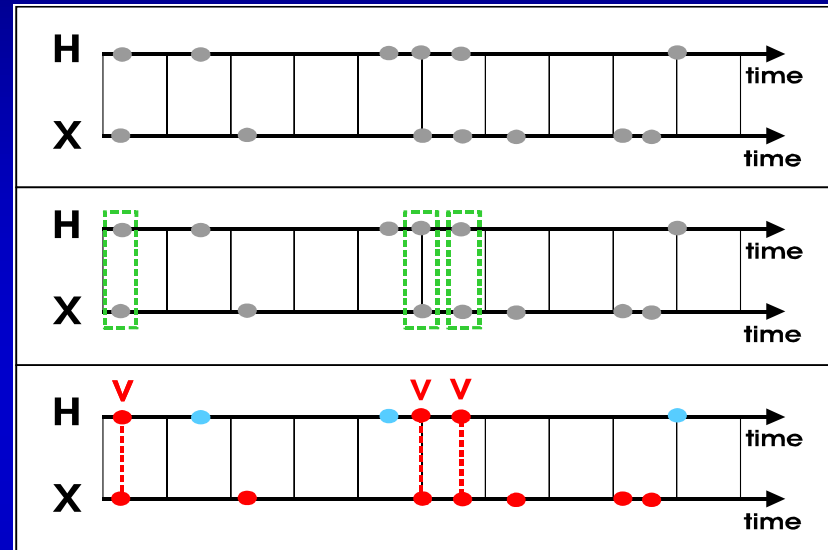
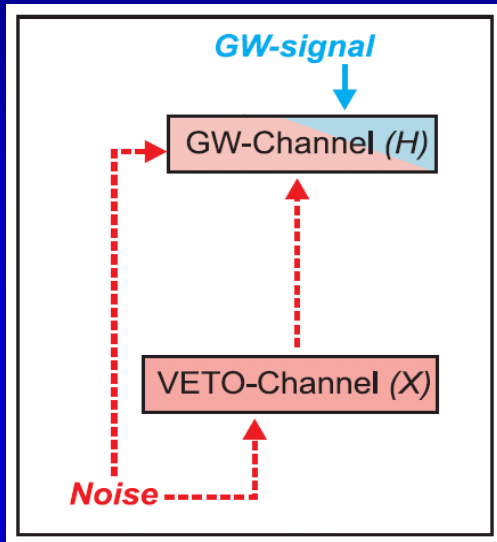




# Standard statistical veto



- Events in H originate from GW and Noise (recorded in X)



- Events in H that occur at the same time as events in X are vetoed.

$$| t_0^H [i] - t_0^X [j] | < t_{\text{win}}$$

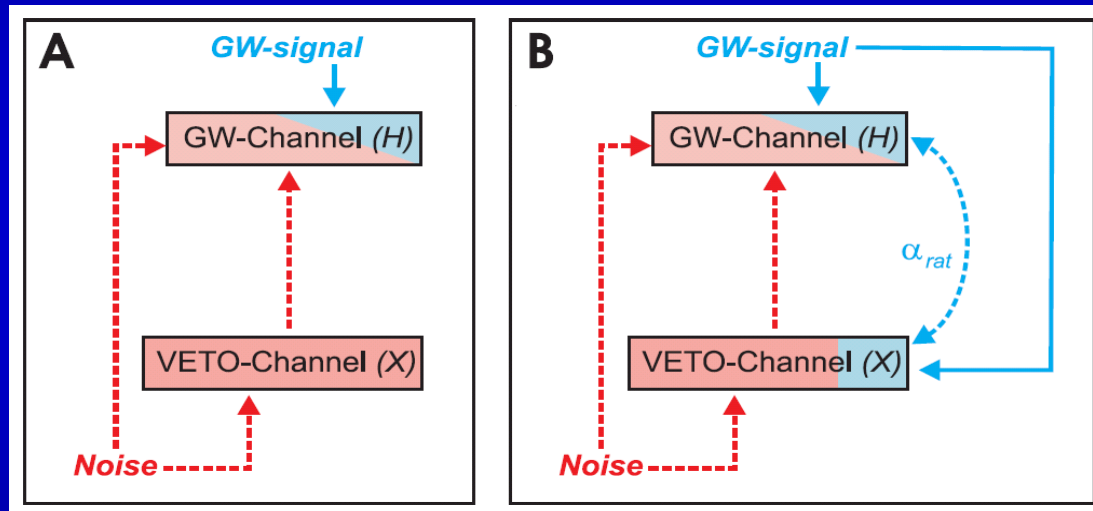
- The standard statistical veto only works for veto channels containing no traces of GW signal (seismometers, microphones, magnetic field sensors, ...).



# Limitations of the standard statistical veto



- As soon as X contains GW signals the application of a standard statistical veto would veto potentially real GW signals.
- Unfortunately many promising veto channels may contain traces of GW-signal, for example interferometer signals (light powers, control signals, ...)



Two populations of coincident events:

Events originating from noise

(we want to veto)

GW-like events

(we DON'T want to veto)



# Separate two populations by the amplitude ratio of the coincident events



If event  $X(j)$  originates from the event  $H(i)$  their amplitude ratio has to correspond to:

$$\frac{a^X[j]}{a^H[i]} = |\alpha_{\text{rat}}[i]|$$

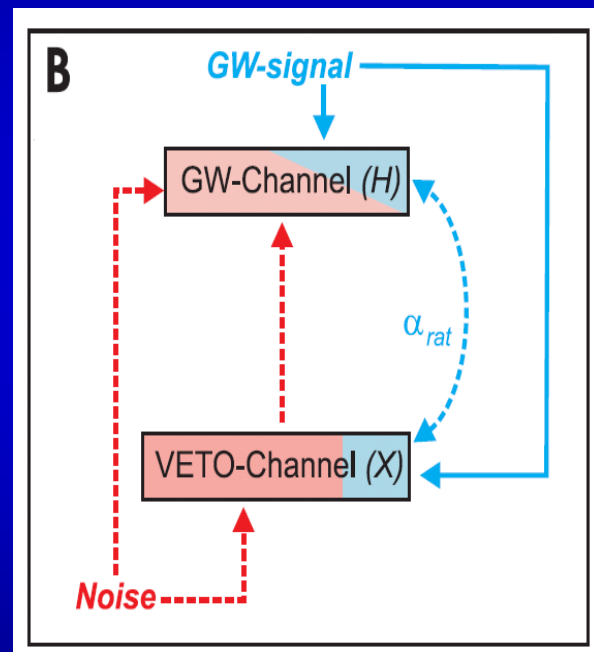
To get a safe veto method we have to compare the amplitude ratio of the two coincident events with the amplitude ratio a GW-signal would have:

If 
$$\frac{a^X[j]}{a^H[i]} = |\alpha_{\text{rat}}[i]|$$

**$H(i)$  is not vetoed**

If 
$$\frac{a^X[j]}{a^H[i]} \neq |\alpha_{\text{rat}}[i]|$$

**$H(i)$  gets vetoed !**





In reality we have to allow for some inaccuracies:

- Error in the amplitude estimation of the two events

$$\Delta a^H [i] \quad \Delta a^X [j]$$

- Error in back-coupling transfer function (measurement, non stationarity)

$$|\alpha_{\text{rat}} [i]|$$

Allow for overall error  $\Delta a_{\text{tot}}$

## VETO CONDITION

Two coincident events  $H(i)$  and  $X(j)$  are vetoed in the case that the amplitude ratio matches one of these requirements:

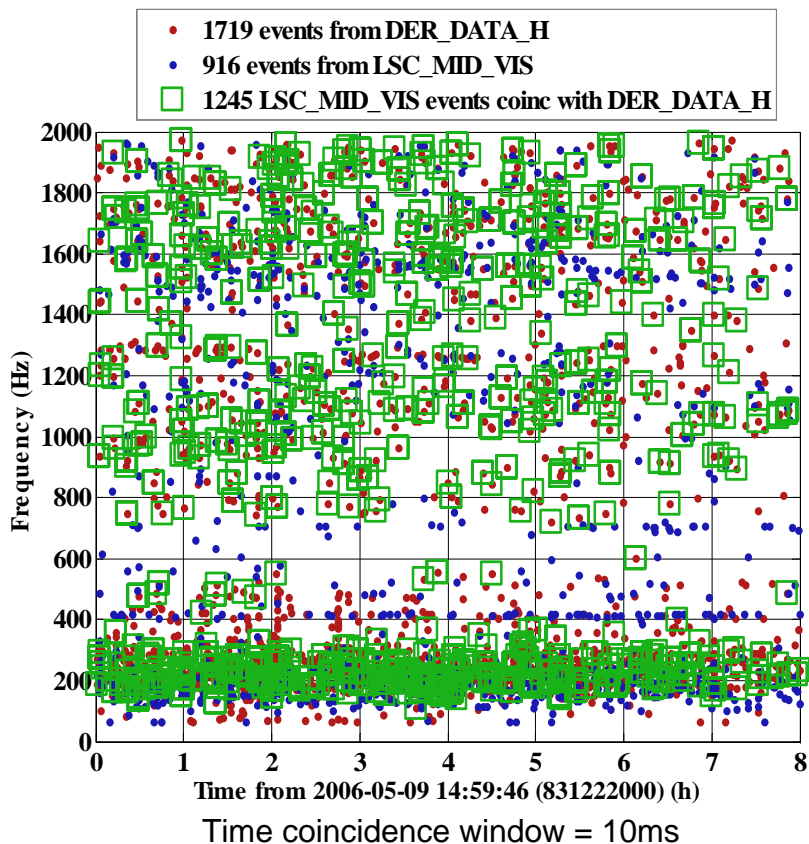
$$\frac{a^X [j]}{a^H [i]} < \frac{|\alpha_{\text{rat}} [i]|}{(1 + \Delta a_{\text{tot}})},$$

$$\frac{a^X [j]}{a^H [i]} > |\alpha_{\text{rat}} [i]| (1 + \Delta a_{\text{tot}})$$

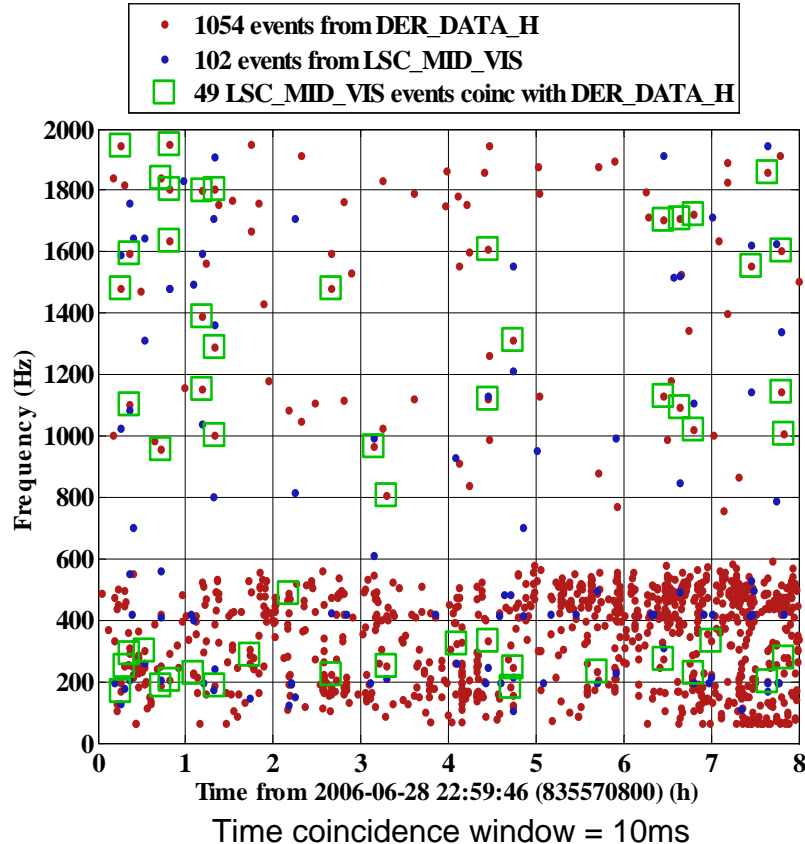
# Dust falling through main output beam



## high dust concentration (broken AC)



## low dust concentration



When dust is falling through the main output beam, coincidence glitches are induced to  $H$  and  $P_{DC}$ .



# $P_{DC}$ contains traces of GW-signal

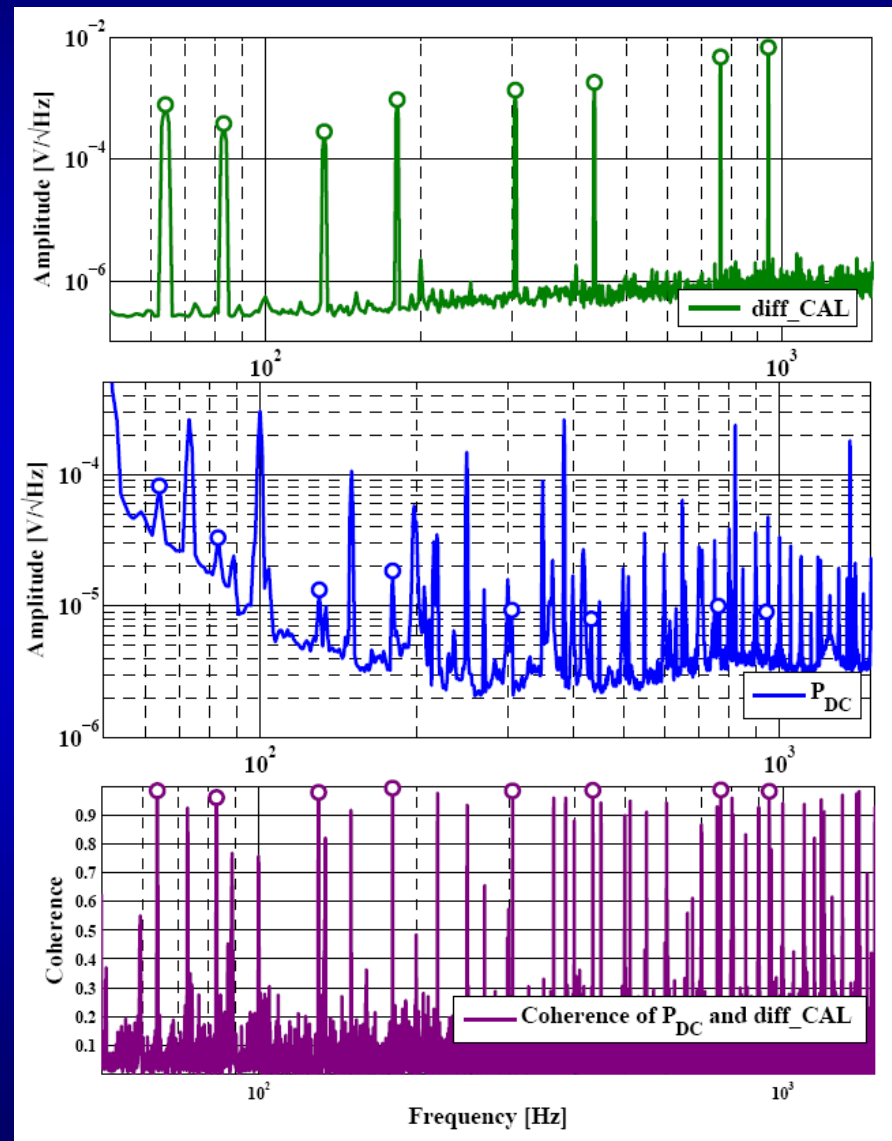


What is  $P_{DC}$  ?

It is the DC light from the main dark port photo detector.

It contains traces of GW-signal.

Hardware injections of sinusoidal signals show coherence of 1.





# Application of the method (Example: ‚Dust-Veto‘)



Application to two data sets of GEO S5 data:

- Data Set 1: Full September 2006 (low dust concentration)
- Data Set 2: 8 hours from May 2006 (high dust concentration)

## Final set of three veto conditions:

$$|t_0^X[j] - t_0^H[i]| < 8 \text{ ms}$$

**Time coincidence**

$$|f_0^X[j] - f_0^H[i]| < 1 \text{ kHz}$$

**Frequency coincidence**

$$\frac{a^X[j]}{a^H[i]} > 3 \alpha_{\text{rat}}[i]$$

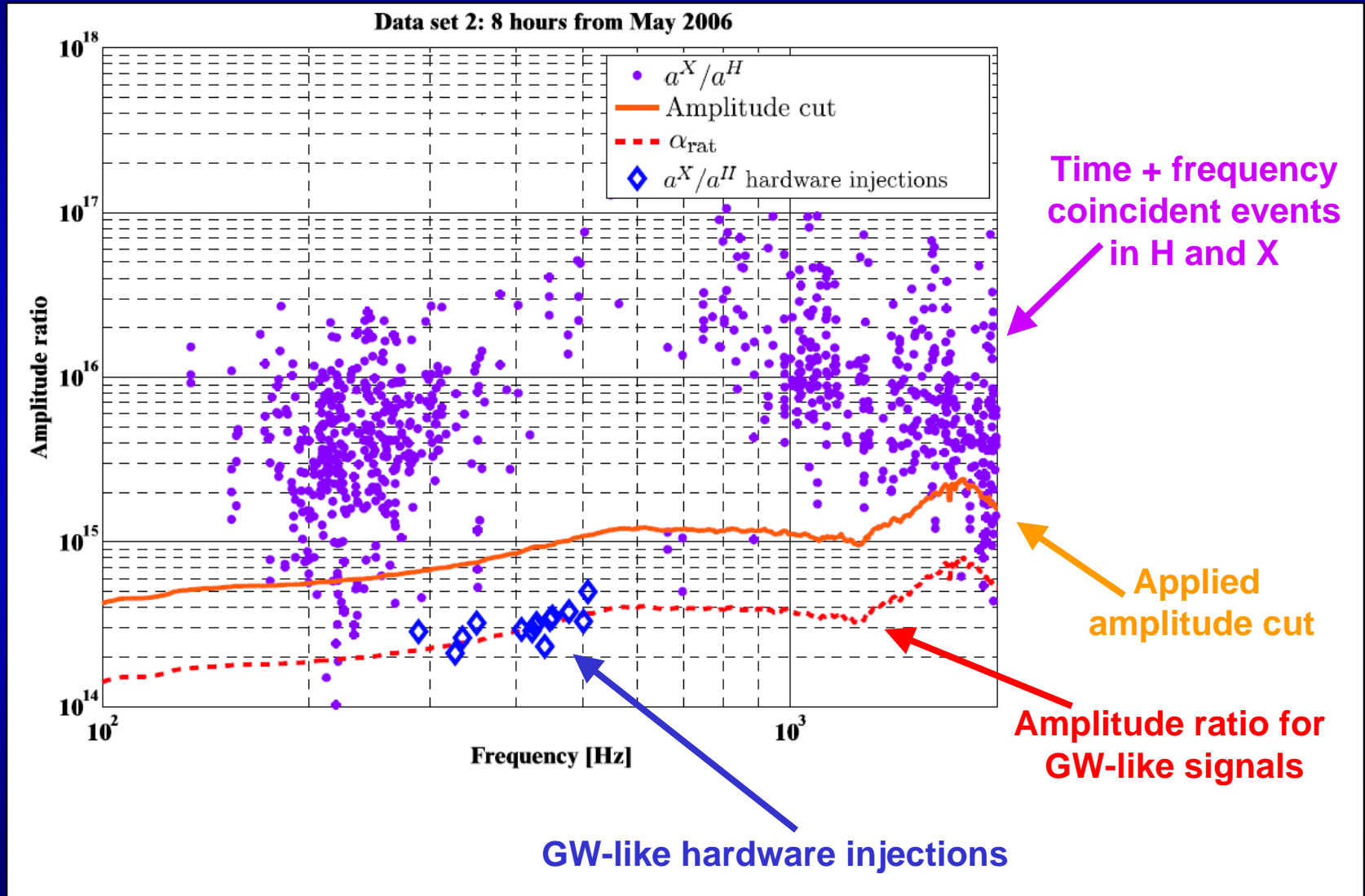
**Amplitude cut**

(amplitude consistency check)

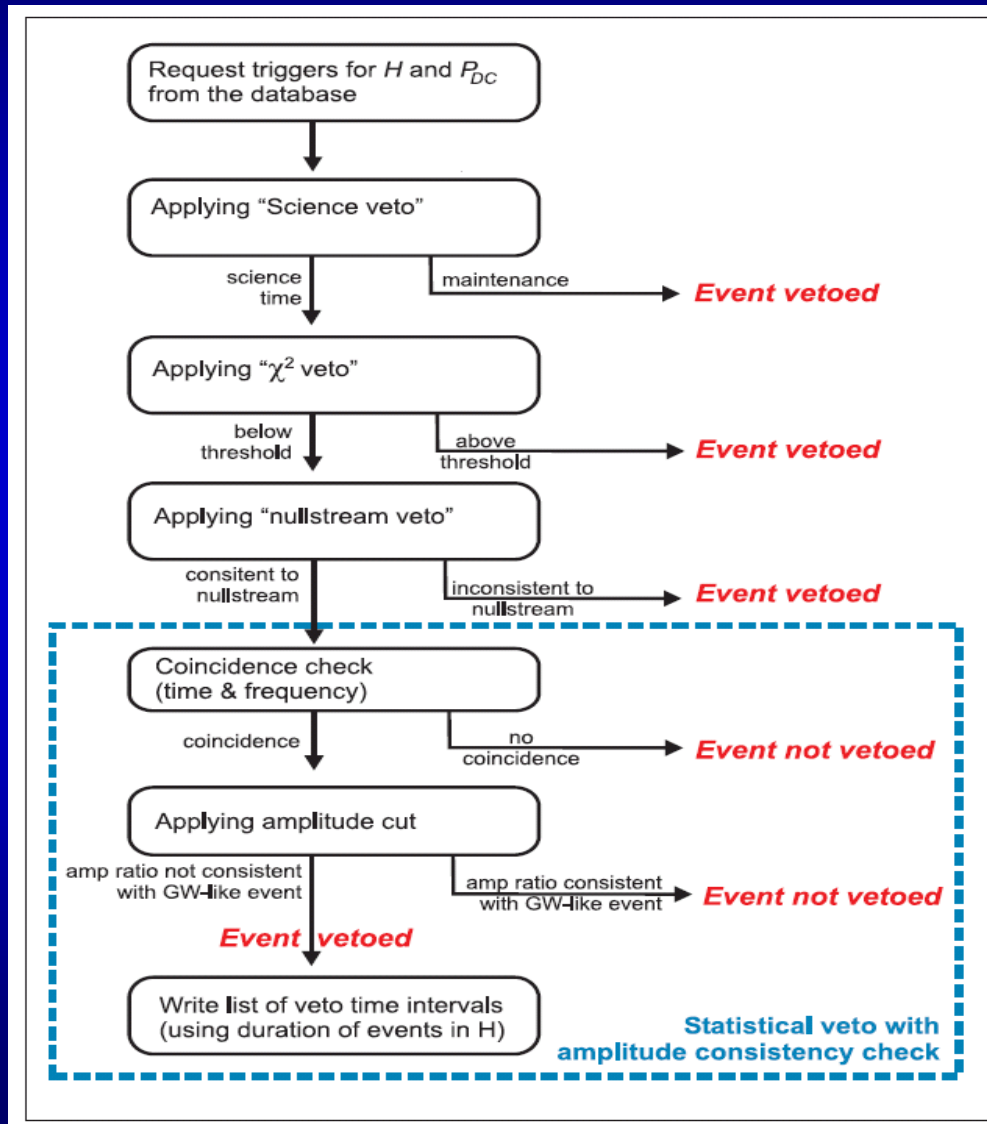




# Dust-Veto: High dust concentration period (Data set 2)



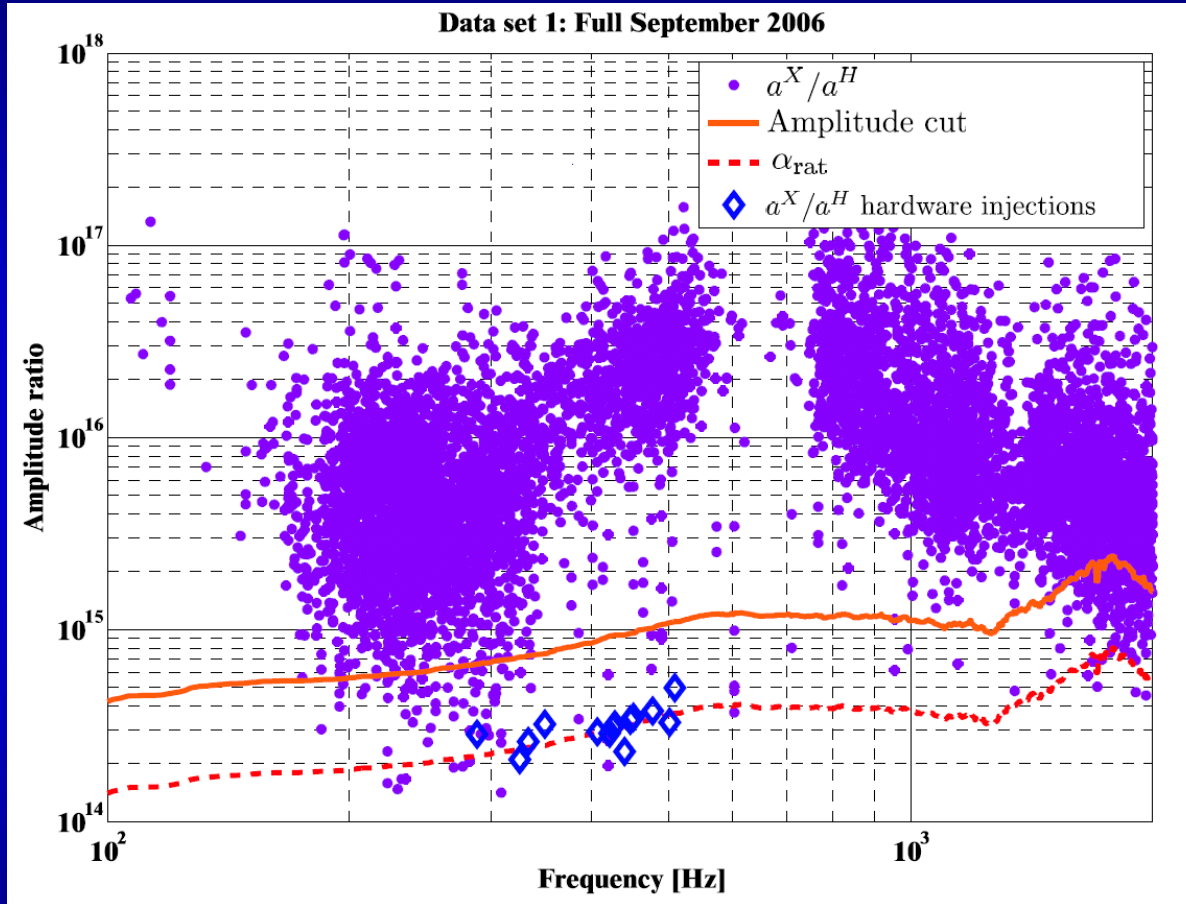
# Full veto pipeline (for GEO S5 data)





# Dust-Veto:

## Low dust concentration period (Data set 1)



Application to S5 data from GEO600 gives encouraging results.



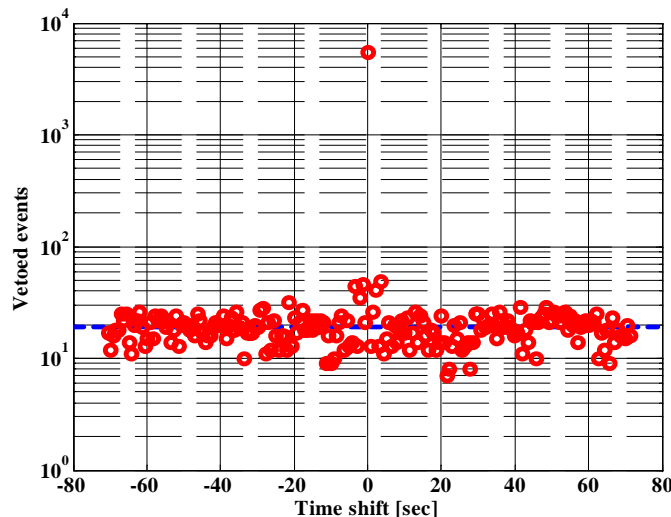
# Summary of the veto performance



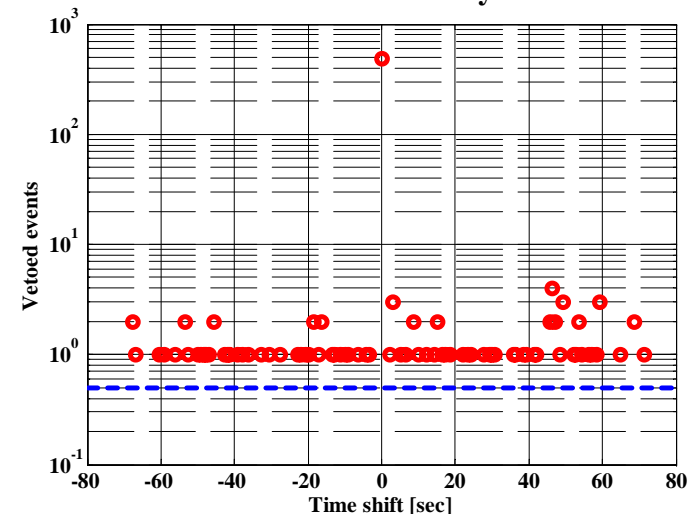
Data Set	1	2
Total number of events in $H$	96454	2281
Total number of events in $\mathcal{P}_{\text{DC}}$	26600	615
Event rate in $H[\text{h}^{-1}]$	134	285
Event rate in $\mathcal{P}_{\text{DC}}[\text{h}^{-1}]$	37	77
Number of events vetoed	5517	491
Efficiency [%]	5.72	21.5
Background [%]	0.02	0.02
Significance	286	1075
Use-percentage [%]	20.7	79.8

*S. Hild et al: „A statistical veto employing an amplitude consistency check “ , submitted to Class. Quantum Grav.*

**Data set 1: Full September 2006**

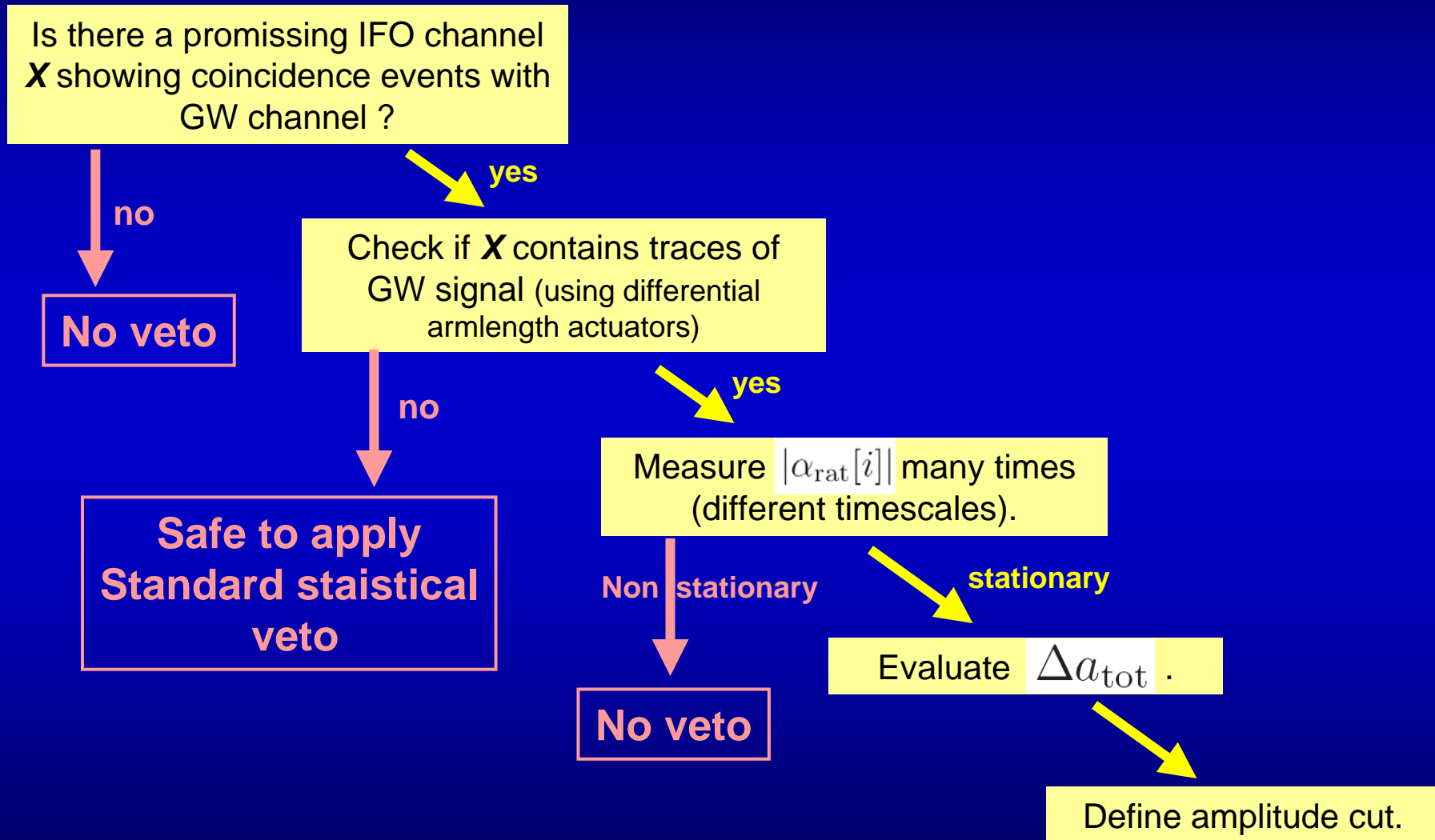


**Data set 2: 8 hours of May 2006**



This new method is easily applicable for all other GW detectors.

# Short recipe for statistical veto with amplitude consistency check





END

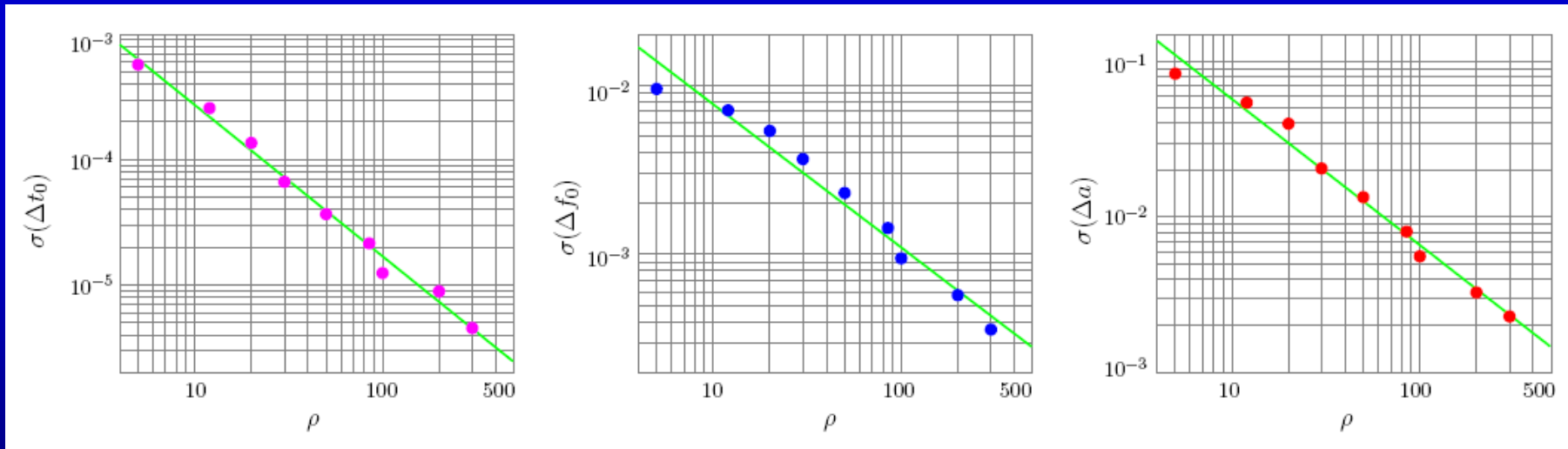


# Determine overall error



Need to determine  $\Delta a_{\text{tot}}$  !!

1. Back-coupling TF was measured to vary less than +/-50% over months.
2. Maximum error in amplitude estimation of mHACR using 3 sigma gives 60% for events of SNR = 4  
(sine-Gaussian injections into Gaussian noise)



3. For the real data we will allow for 200% error in amplitude estimation.