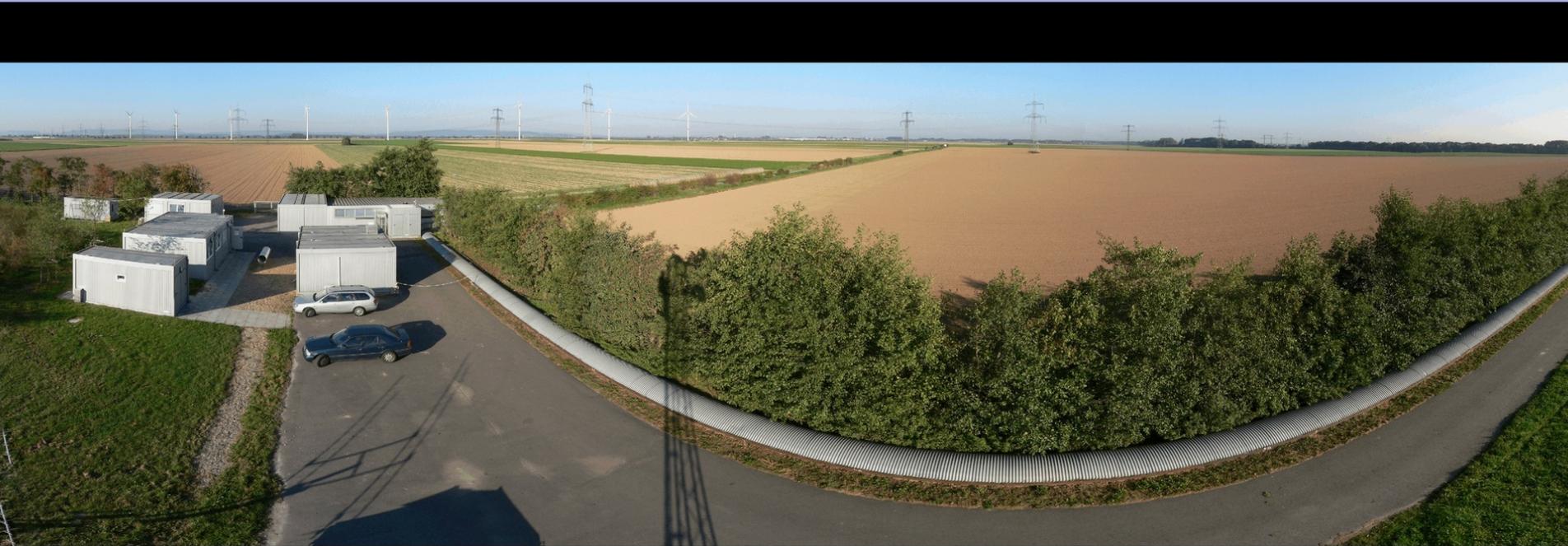


Comparison of tuned and detuned Signal-Recycling



Stefan Hild for the GEO-team

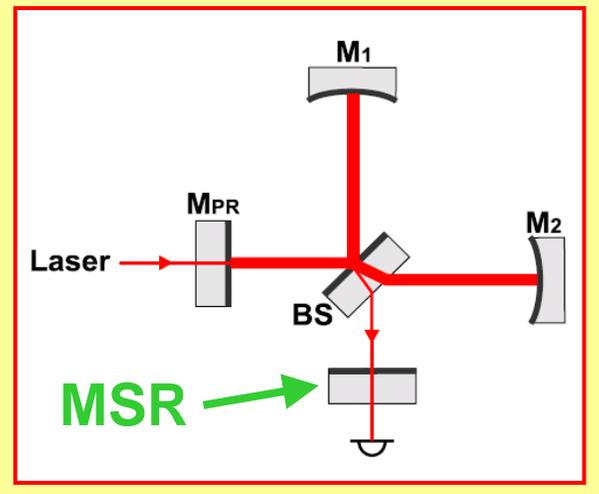




Signal-Recycling in short

An additional recycling mirror
(MSR) at the dark port allows:

- enhancing the GW signal
- shaping the detector response



Two main parameters:

- **Bandwidth** (of the SR resonance)

broadband

narrowband

- **Tuning** (Fourier frequency of the SR resonance)

tuned

detuned

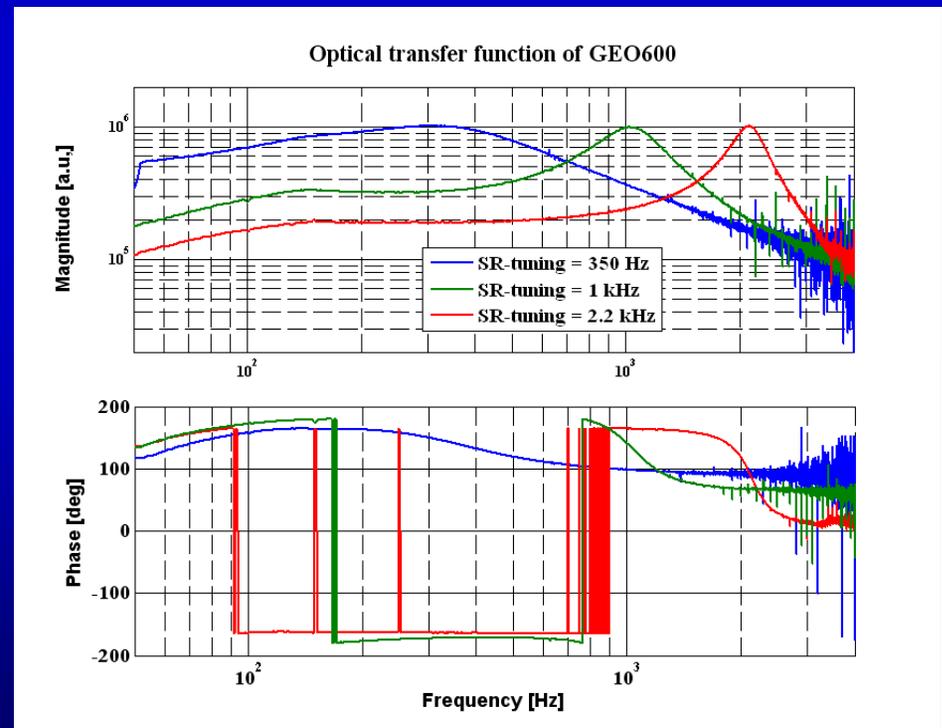


SR-Tuning in GEO so far

- For various reason we are not able aquire lock for tuned SR.
- Lock takes place at a detuning of a few kHz.
- Afterwards the aquisition detector is tuned to ist operation point.
- So far tunings between 5kHz and 250 Hz had been realized.

Downtuning in steps of 25 Hz, 6 Parameters need to be adjusted:

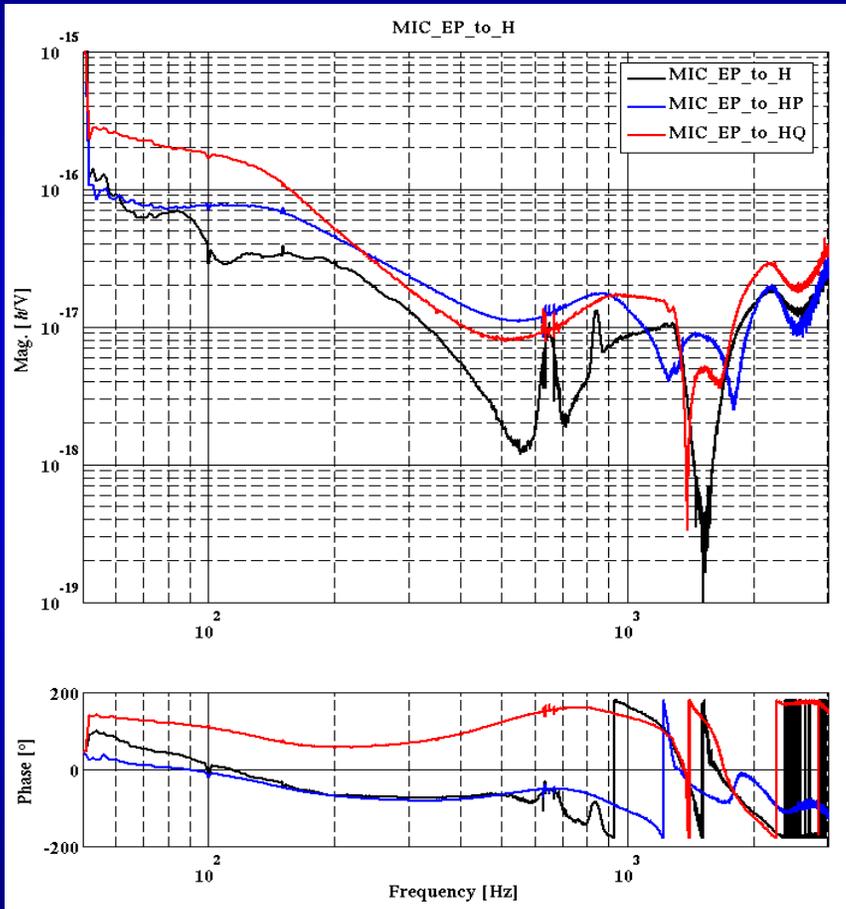
- SR frequency
- SR gain
- SR phase
- MI gain
- MI phase
- MI autoalignment gain



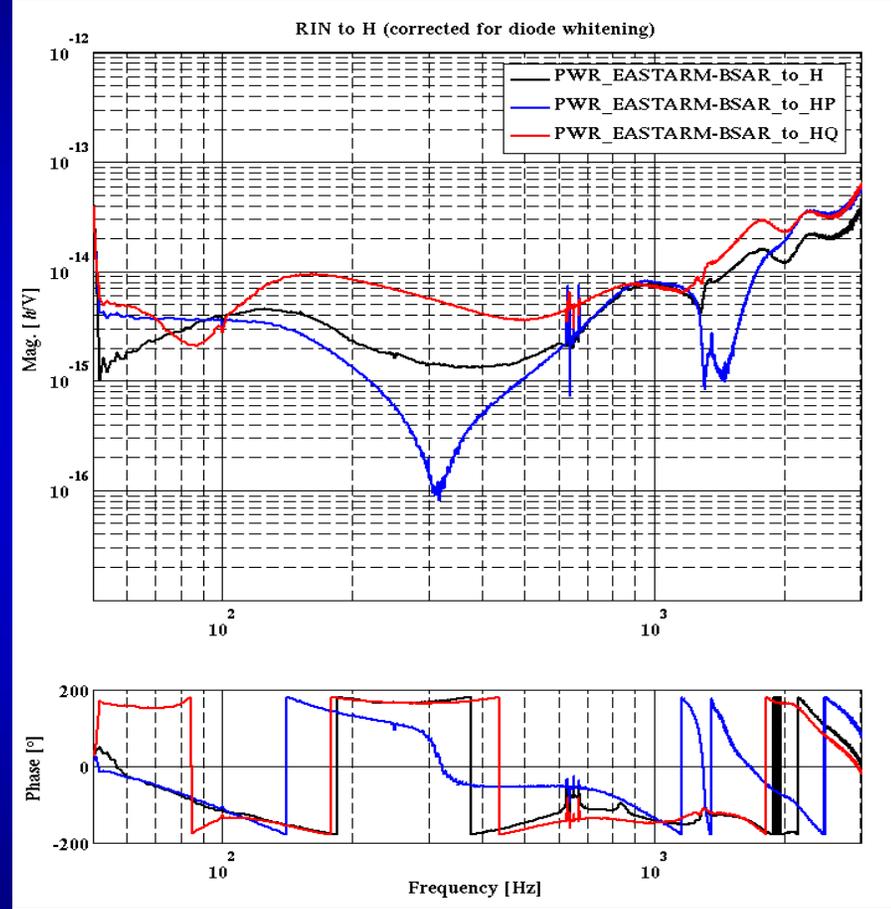


Motivation (1): Detuned SR complicates noise couplings and TFs

Frequency noise coupling to $h(t)$



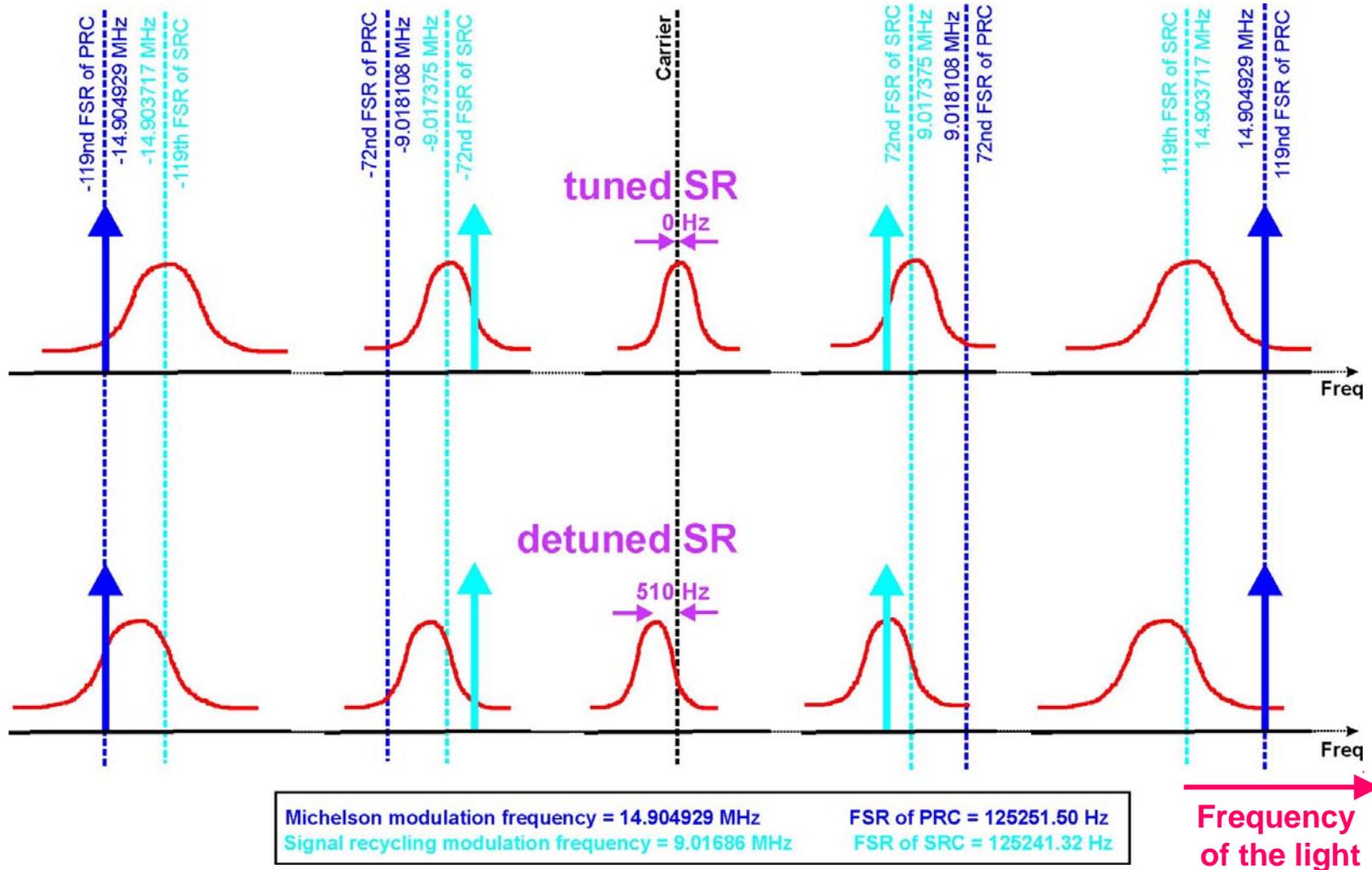
Laser intensity noise coupling to $h(t)$



In a detuned detector TF may become complicated due to interaction and different resonance conditions of various sidebands.

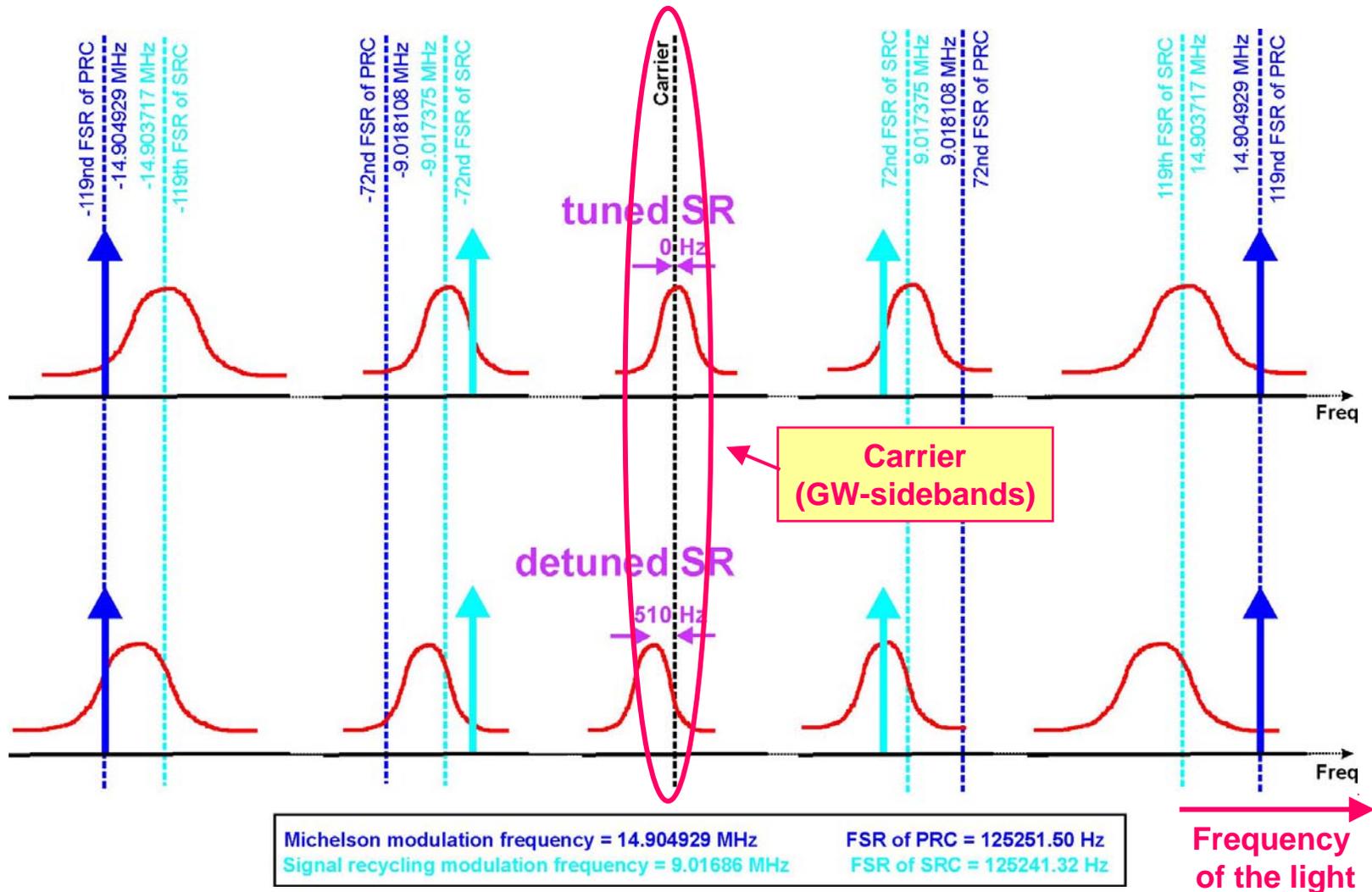


Motivation (2): Sideband picture for tuned and detuned SR



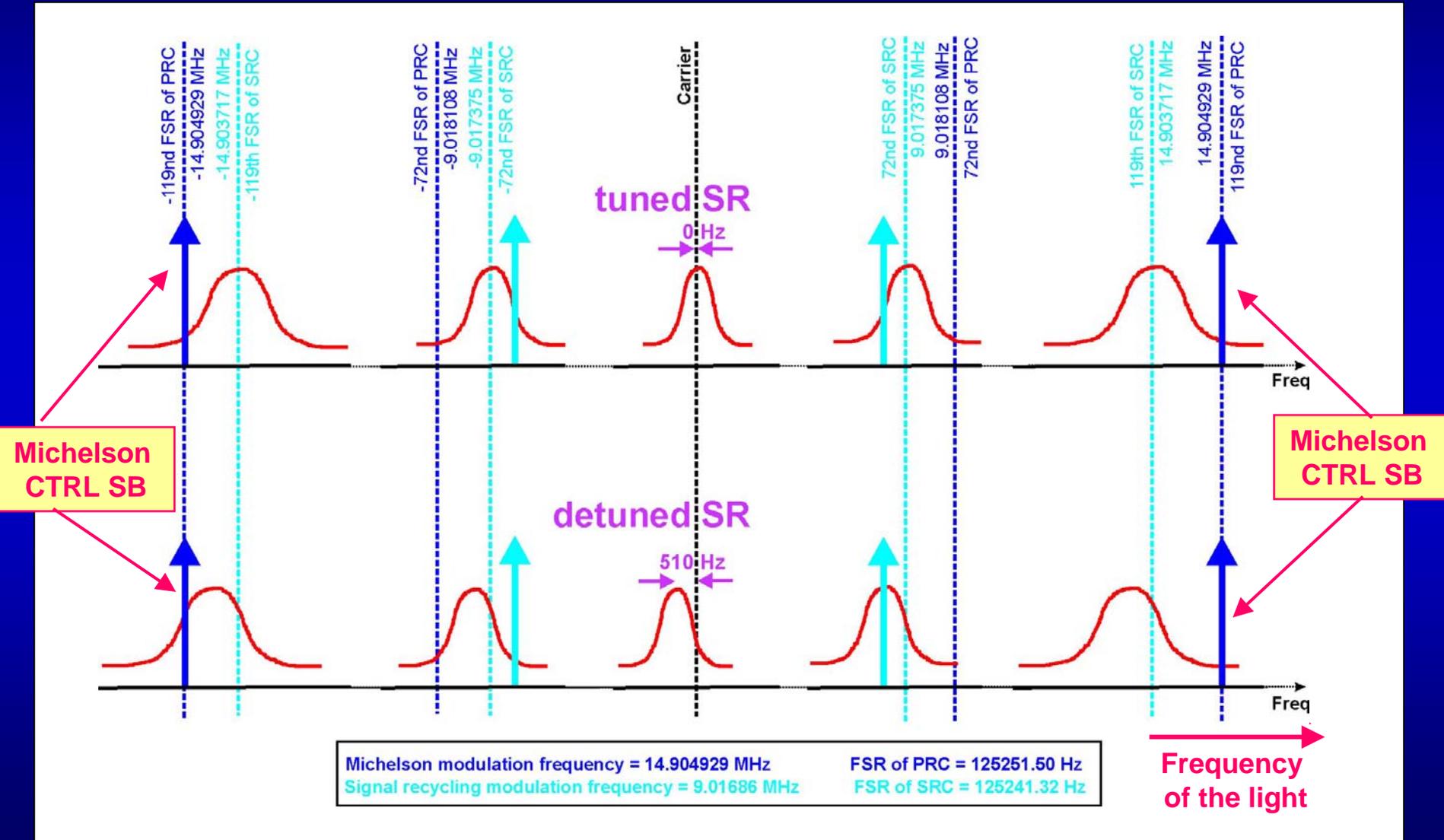


Motivation (2): Sideband picture for tuned and detuned SR



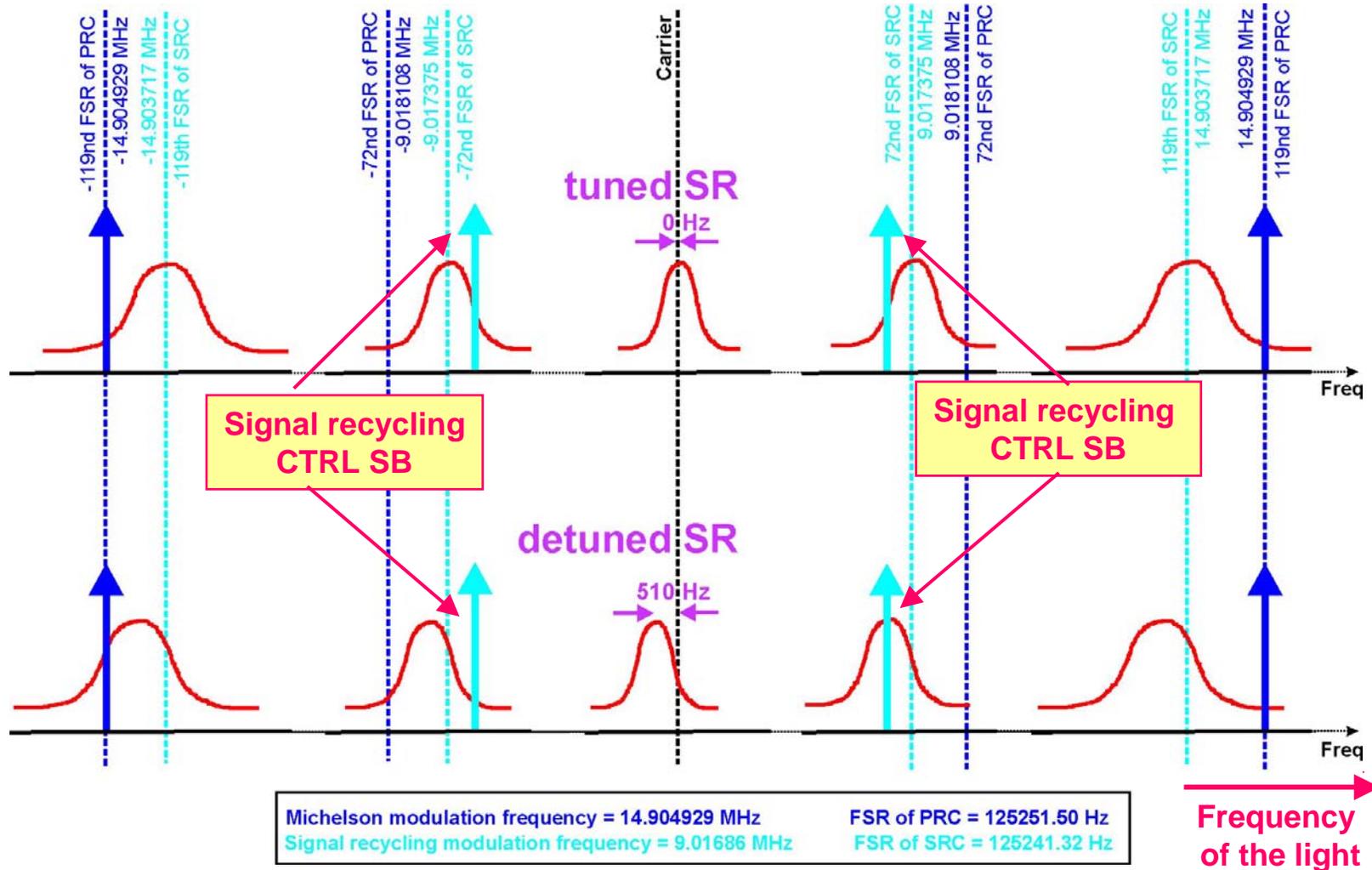


Motivation (2): Sideband picture for tuned and detuned SR



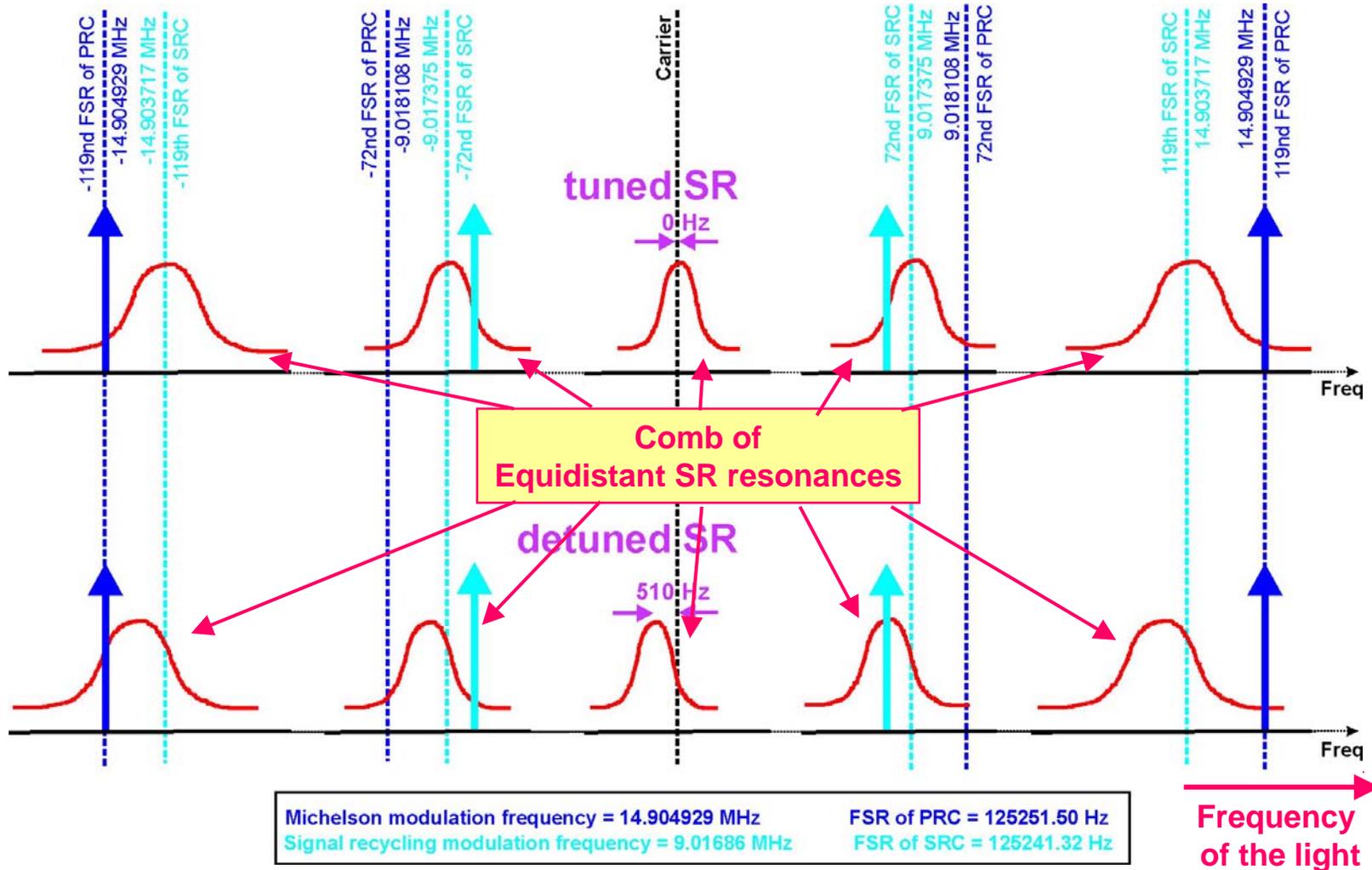


Motivation (2): Sideband picture for tuned and detuned SR



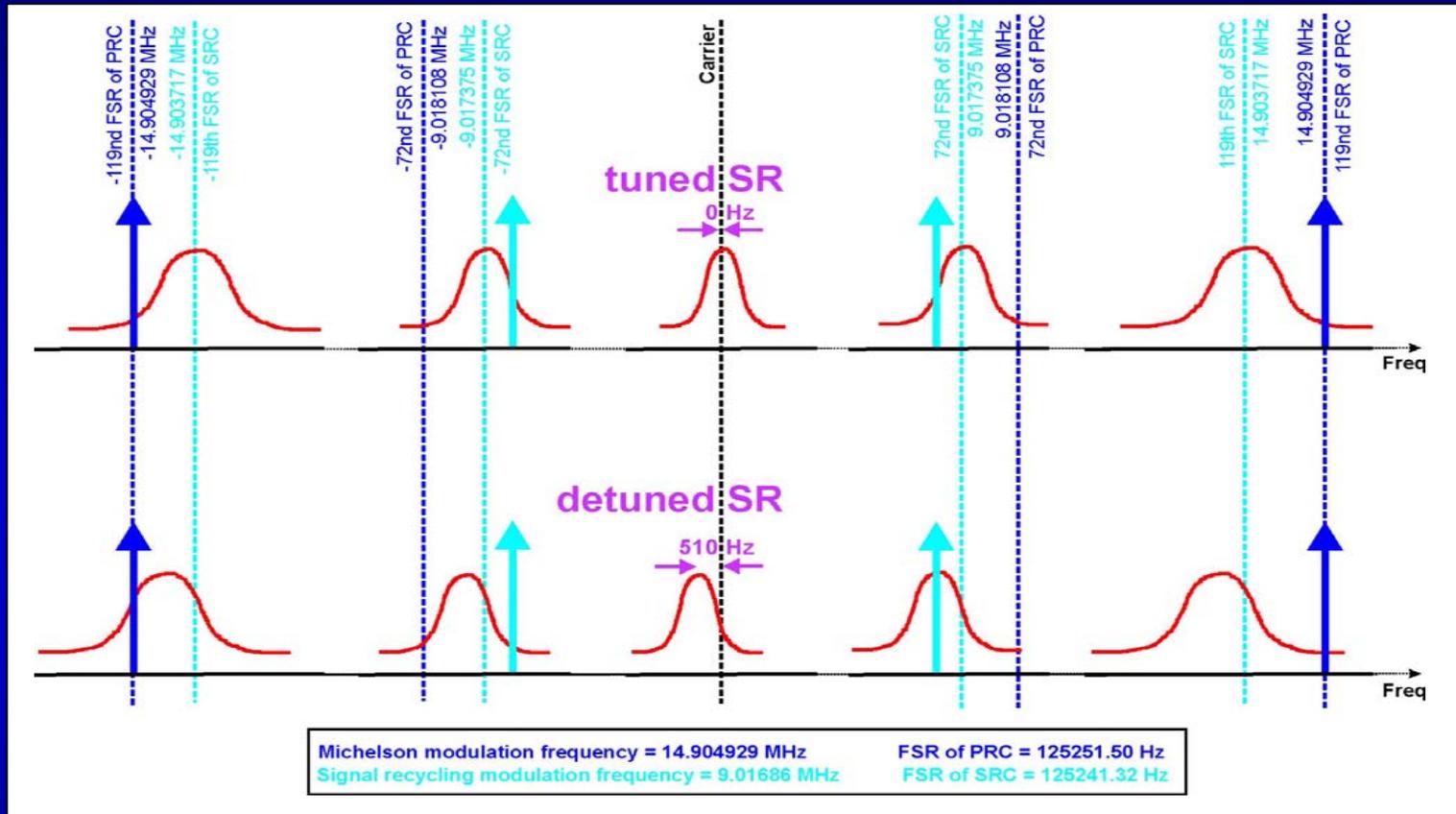


Motivation (2): Sideband picture for tuned and detuned SR





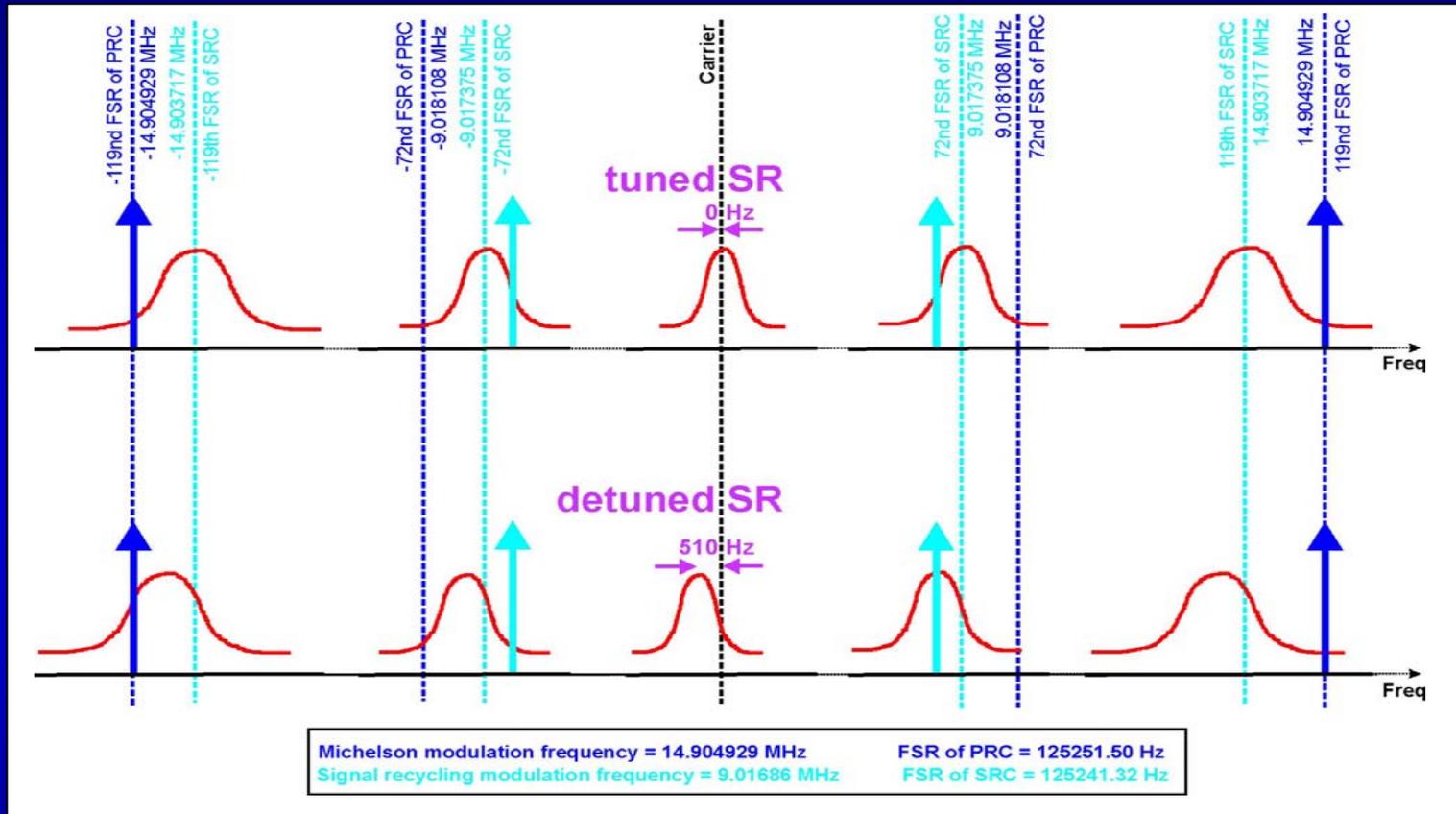
Motivation (3): Sideband picture for detuned SR (510 Hz)



- MSR is locked onto upper SB => each SR resonances is 510 Hz left of a FSR of SRC.
- Upper SR-SB resonant, lower SR-SB off resonance.
- Lower MI-SB nearly resonant, upper MI-SB far off resonance.
- Upper and lower GW signal see different completely different resonance condition.



Motivation (4): Sideband picture for tuned SR



- Going from detuned (510Hz) to tuned: comb of SR resonances is shifted 510 Hz to the right.
- Upper and lower SR-SB see the same resonance condition (nearly resonant).
- Upper and lower MI-SB see the same resonance condition.
- Upper and lower GW signal SB see the same resonance condition.



Motivation for tuned SR

There is the possibility:

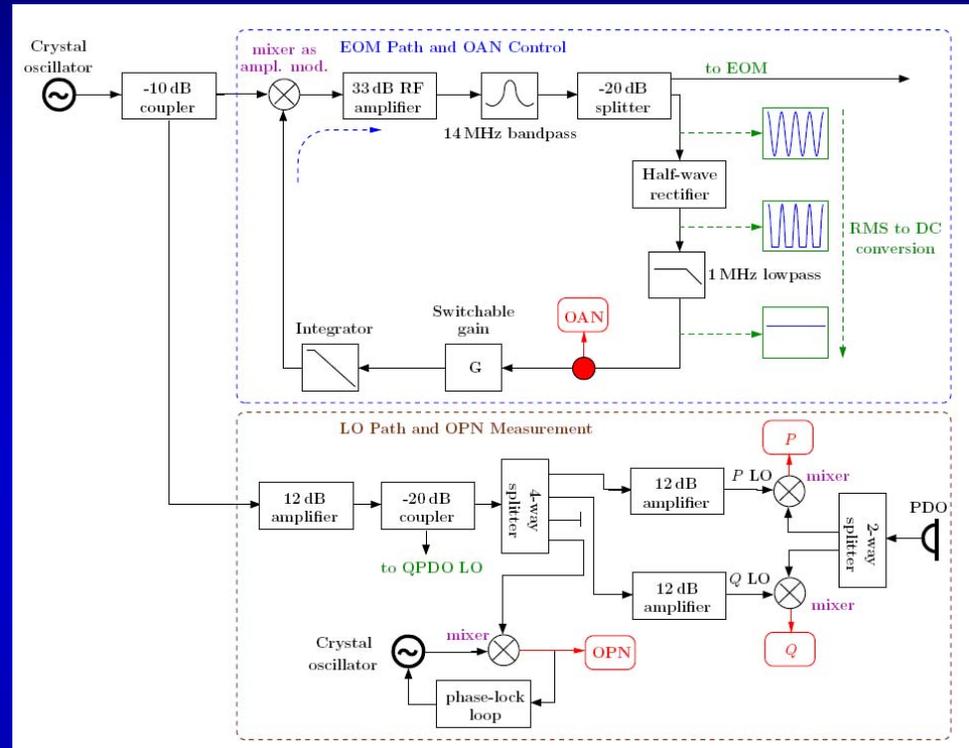
- Get better noise performance due to symmetric sidebands.

Phase noise in P originates from big DC in Q

We believe that a large contribution of the phase noise in P originates from a large DC signal in the Q quadrature.

Therefore it would be nice to reduce the DC signal in Q.

(see Josh's talk)



Here a large DC signal in the Q quadrature mixes OPN of the LO strongly into the P quadrature,

$$\phi_P = \langle Q \rangle \phi_{\text{OPN}}, \quad (3.10)$$

where ϕ_P is the phase noise contribution to P , ϕ_{OPN} is the OPN of the LO signal, and $\langle Q \rangle$ is the time-averaged (DC) signal in Q



Motivation for tuned SR

There is the possibility:

- Get better noise performance due to symmetric sidebands.
- Less RF amplitude modulation on main photodiode (reduction of potential saturation / nonlinear effects)

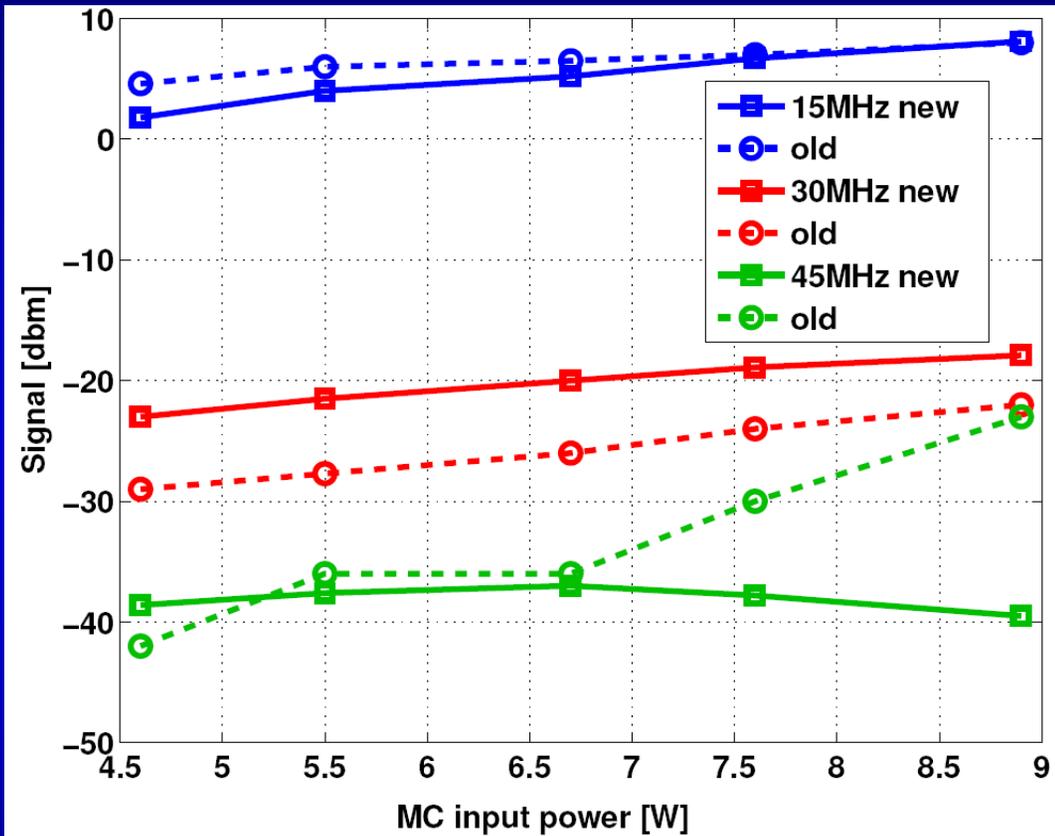
Saturation in our main photo diode

We believe we saw some saturation in the main photodiode due to the large RF amplitude modulation (Q-DC).

The signal voltage (1V @15MHz) modulated the bias voltage (5V) by about 20%.

Problem was for the moment (higher bias, different circuit).

For long term it would be desirable to reduce Q-DC !!



High-power, low-noise, and multiply resonant photodetector for interferometric gravitational wave detectors

Grote H ¹⁸



Motivation for tuned SR

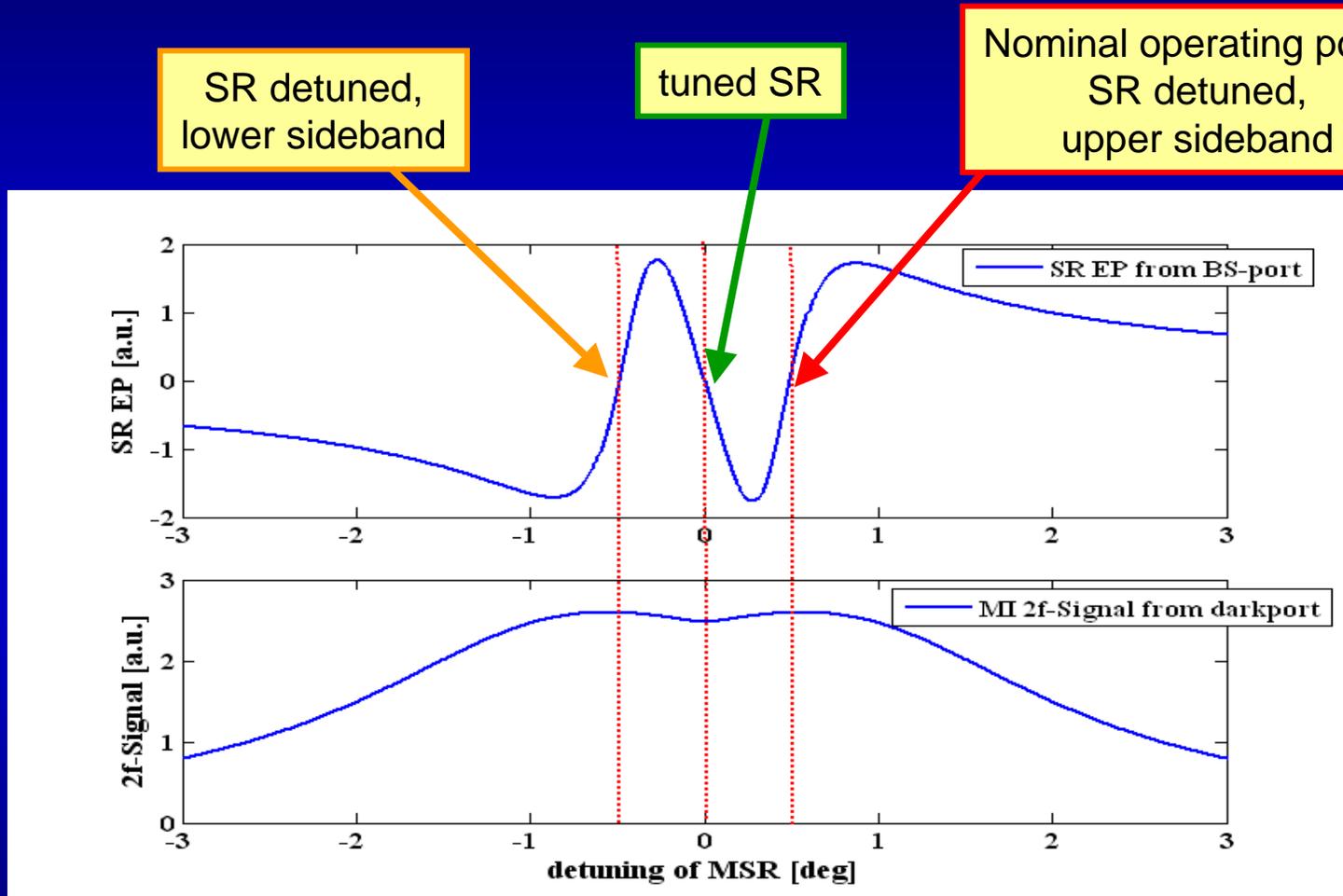
There is the possibility:

- Get better noise performance due to symmetric sidebands.
- Less RF amplitude modulation on main photodiode (reduction of potential saturation / nonlinear effects)
- Get simpler noise couplings and transferfunctions.
- Get a better understanding of the detector.

-
- Nice research / demonstration project



Signal-Recycling errorsignal

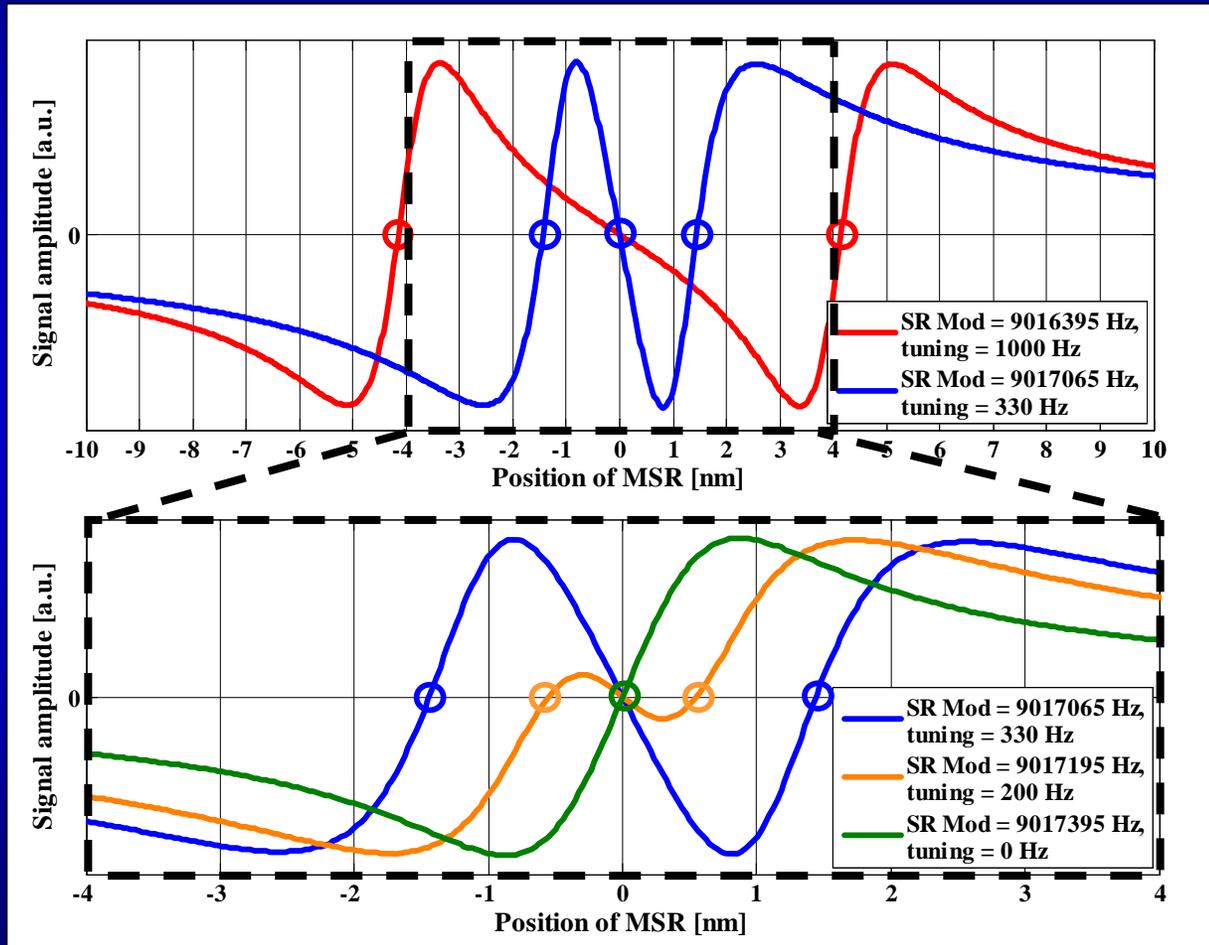


Derived from an RF modulation-demodulation technique.



Why can't we tune down to tuned SR ?

Tuning is done by changing the SR RF-modulation frequency.



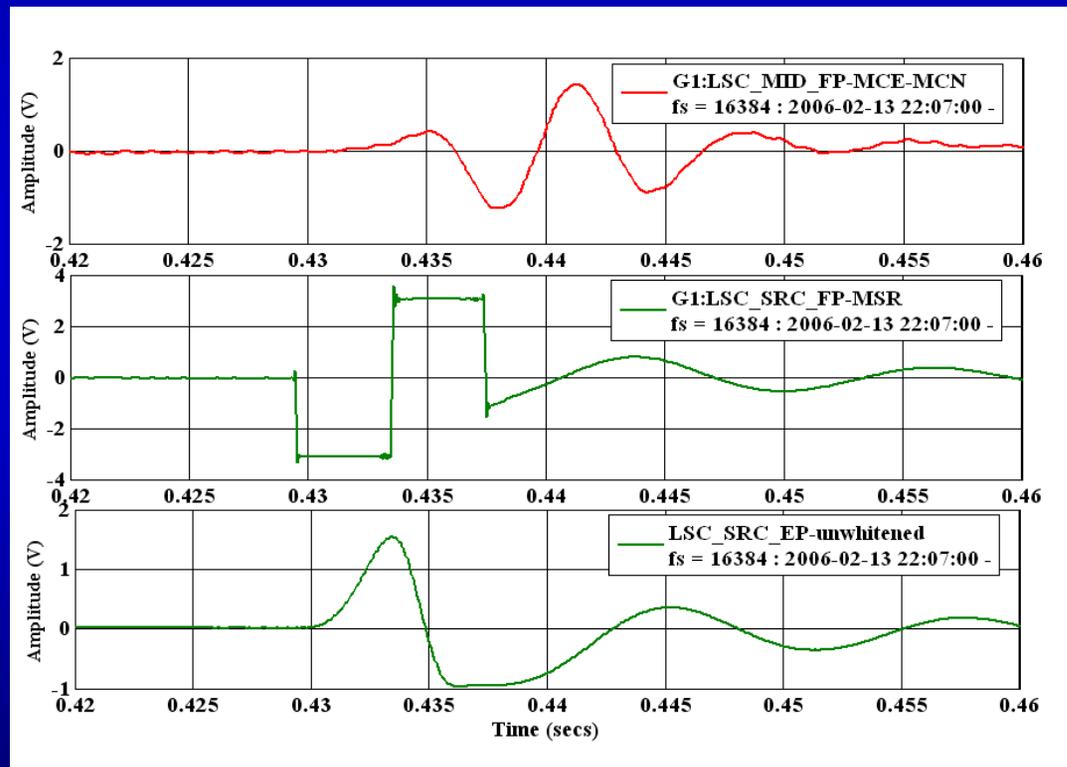
When we tune below 250 Hz the controlsignal vanishes / goes „crazy“.



Kicking MSR

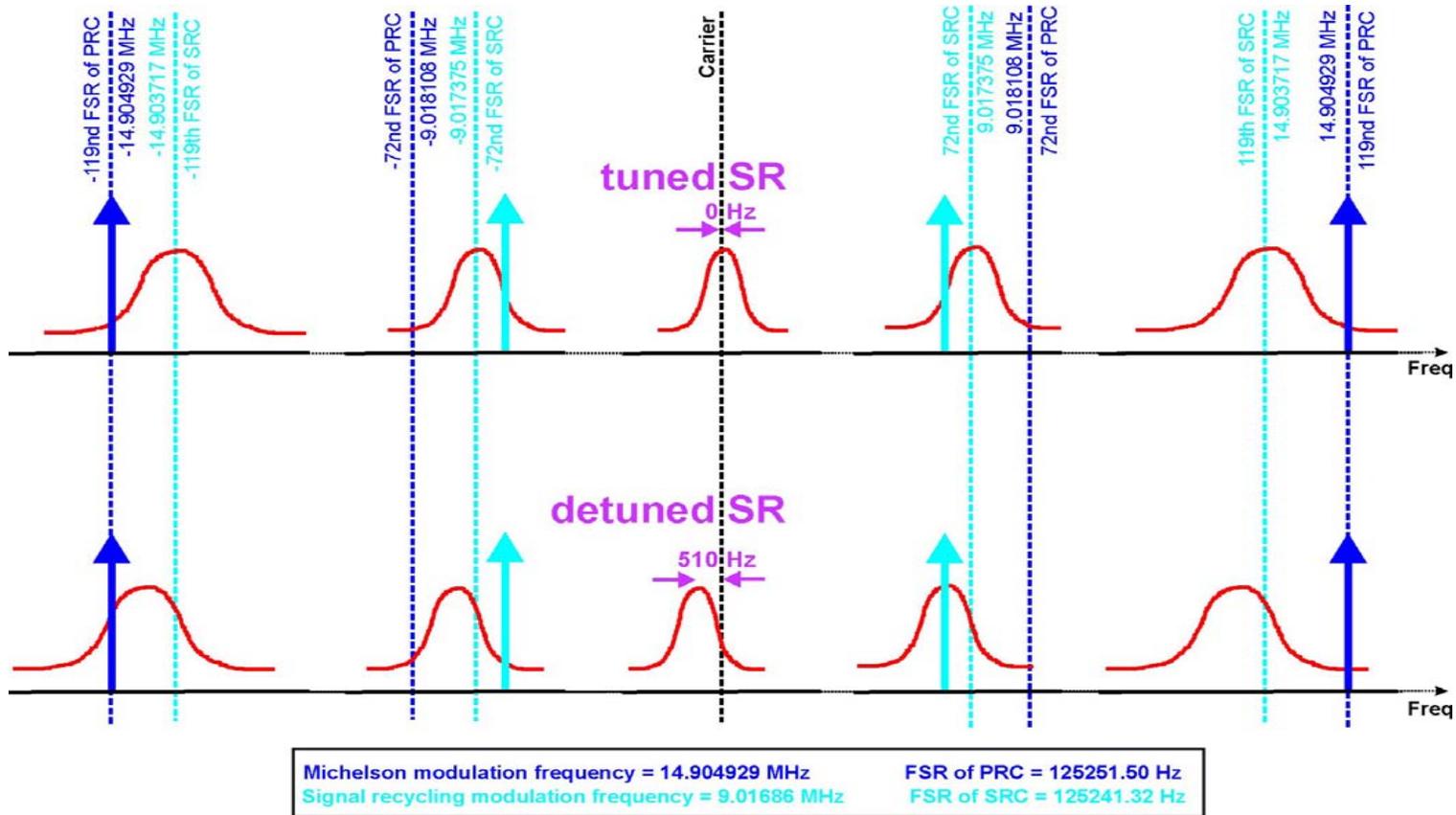
Kicking MSR in a controlled way:

- Fast enough that all other loops can't recognize.
- 4 ms of acceleration and 4 ms of deceleration.



Works fine: Jumping to tuned and to the lower SR sideband

Sideband picture for tuned and detuned SR



From detuned to tuned:
CTRL-signals that are generated from ,carrier‘ need to be adjusted for the different resonance condition of the carrier.

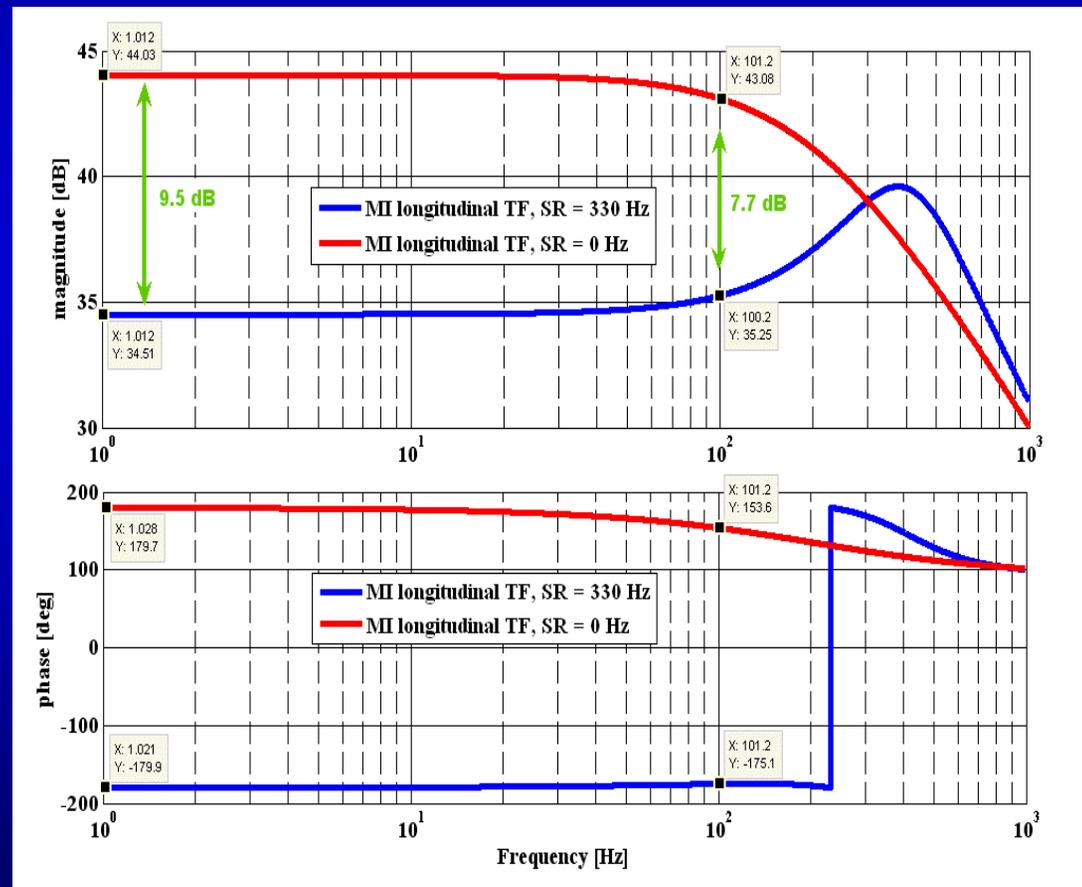


Parameter adjustment

When jumping to tuned SR you need to change a few parameters / compensate for the pole of the SR cavity:

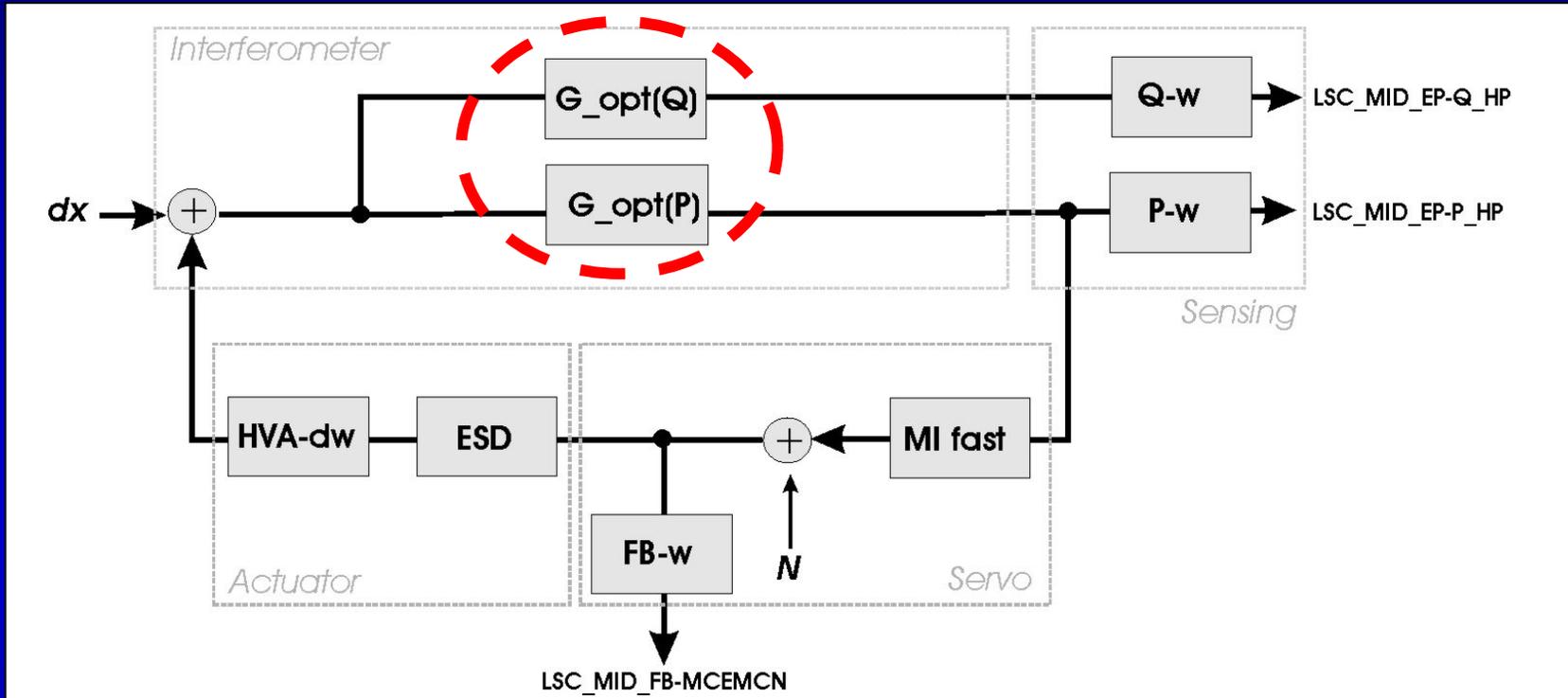
- MI long needs an additional differentiator switched in.
- MIAA gains need to be adjusted.
- Swap sign of SR_EP.
- Adjust SR long gain.

Simulations done with FINESSE





Determine the optical gain for P- and Q-quadrature (TF from diff armlength change to detector output)

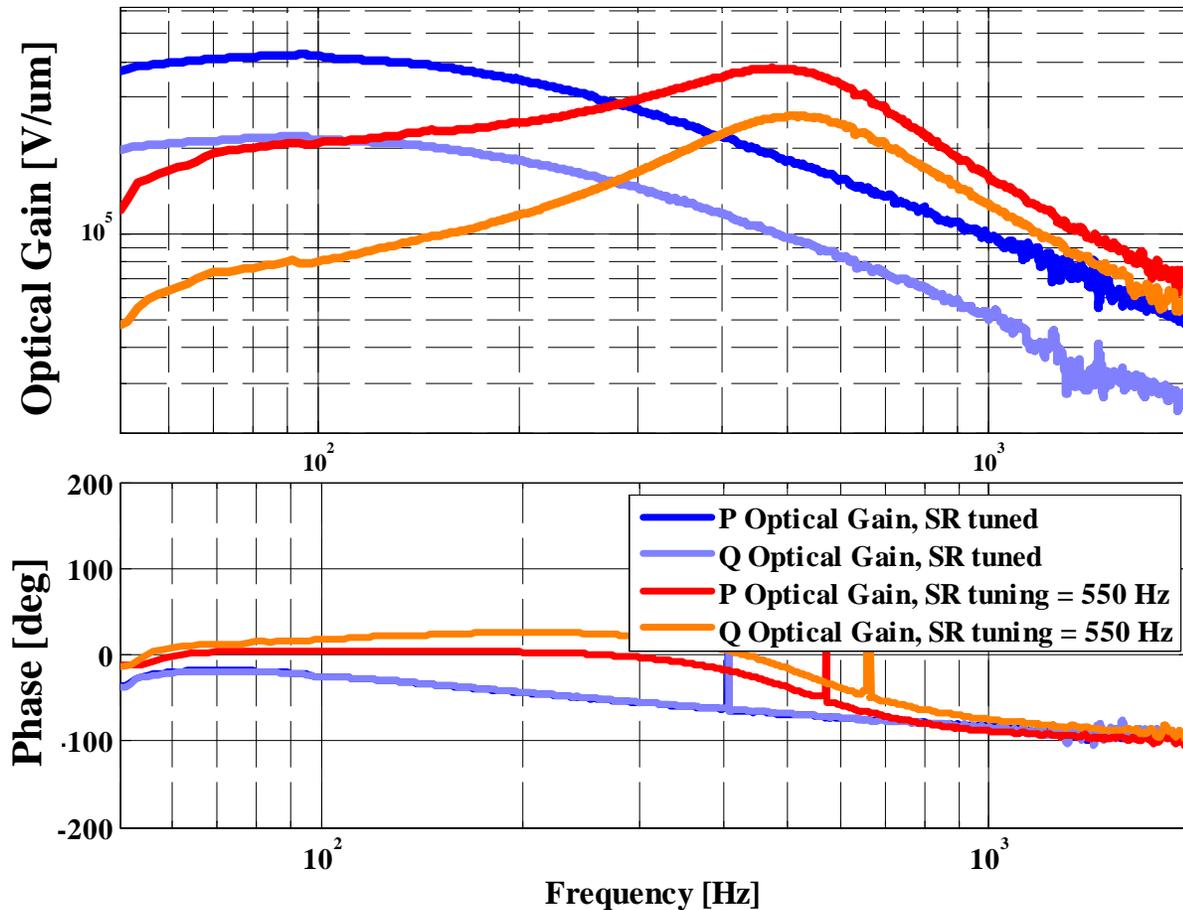


$$\frac{\text{MID_EP-P_HP} \cdot \text{P-dw}}{\text{MID_FB-MCEMCN} \cdot \text{FB-dw}} = \text{ESD} \cdot \text{HWA-dw} \cdot G_{\text{opt}}$$

$$G_{\text{opt}} = \frac{\text{MID_EP-P_HP} \cdot \text{P-dw}}{\text{MID_FB-MCEMCN} \cdot \text{FB-dw}} \cdot \frac{1}{\text{ESD}} \cdot \frac{1}{\text{HWA-dw}}$$



Optical gain for tuned and detuned SR

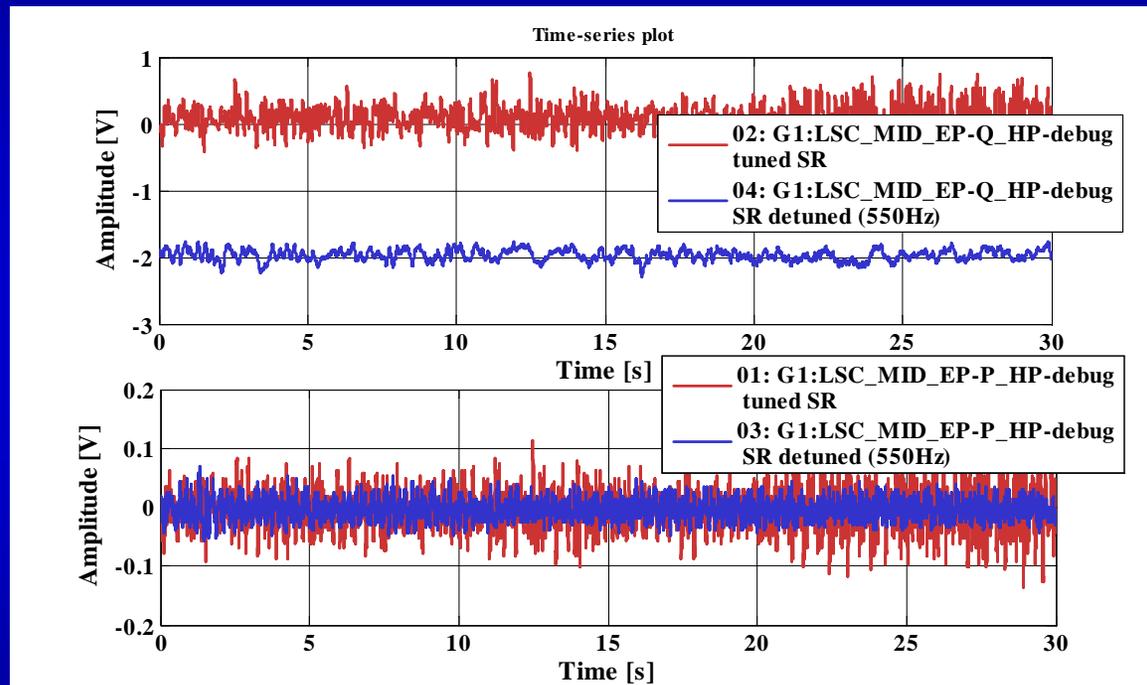


Matches our expectations:
Bandwidth of the TF goes down by a factor of 2.



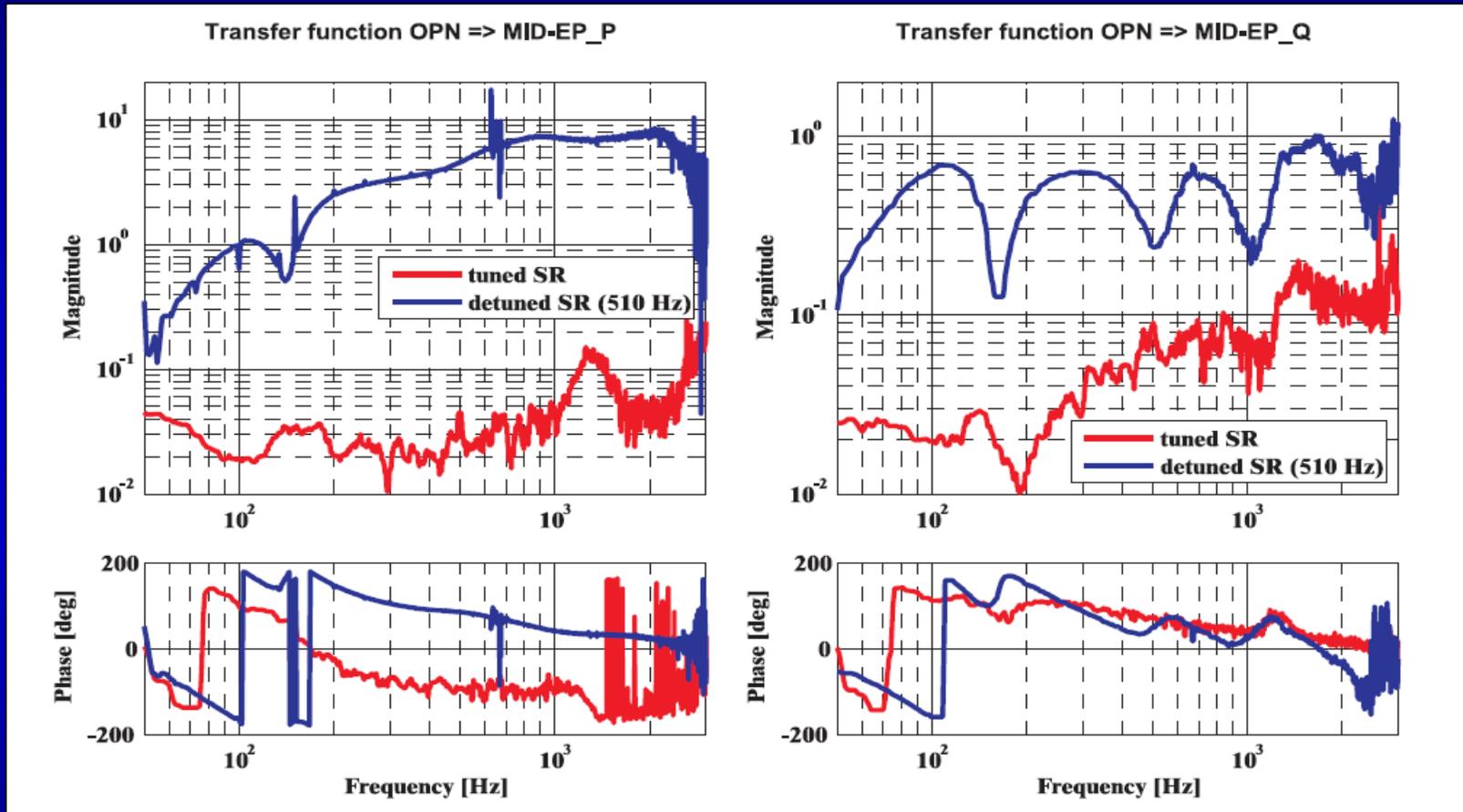
Significantly reduced RF AM.

The signal from the Q-quadrature is in the tuned case reduced to nearly zero (red). RMS of Q-signal is reduced by a factor of 12 !!



- Nice for the photodiode: Less potential saturation / nonlinearities !!
- What is about phase noise in P-quadrature??

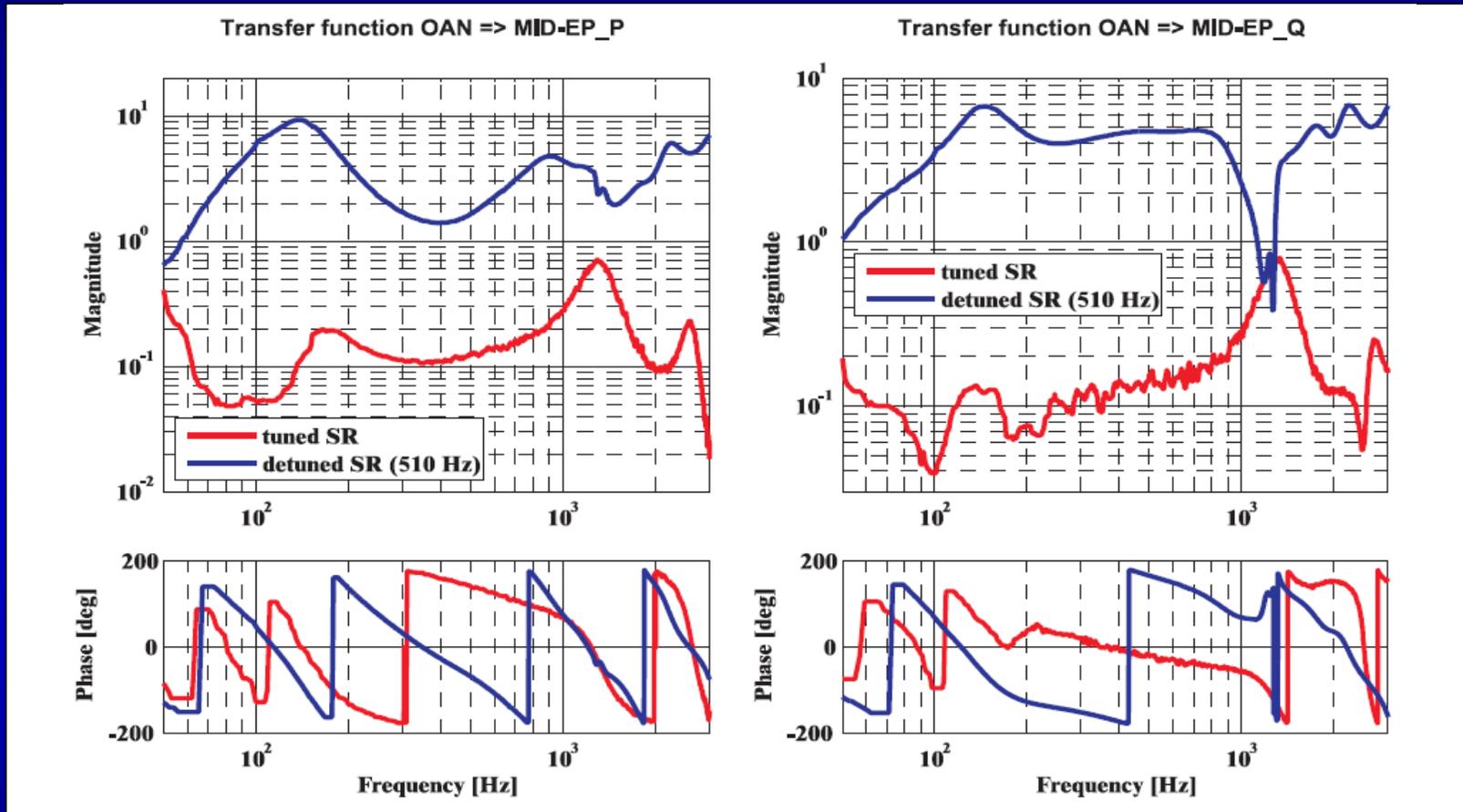
MI Oscillator phase noise (OPN)



**OPN coupling @ tuned SR:
P dramatically reduced, Q significantly reduced.**



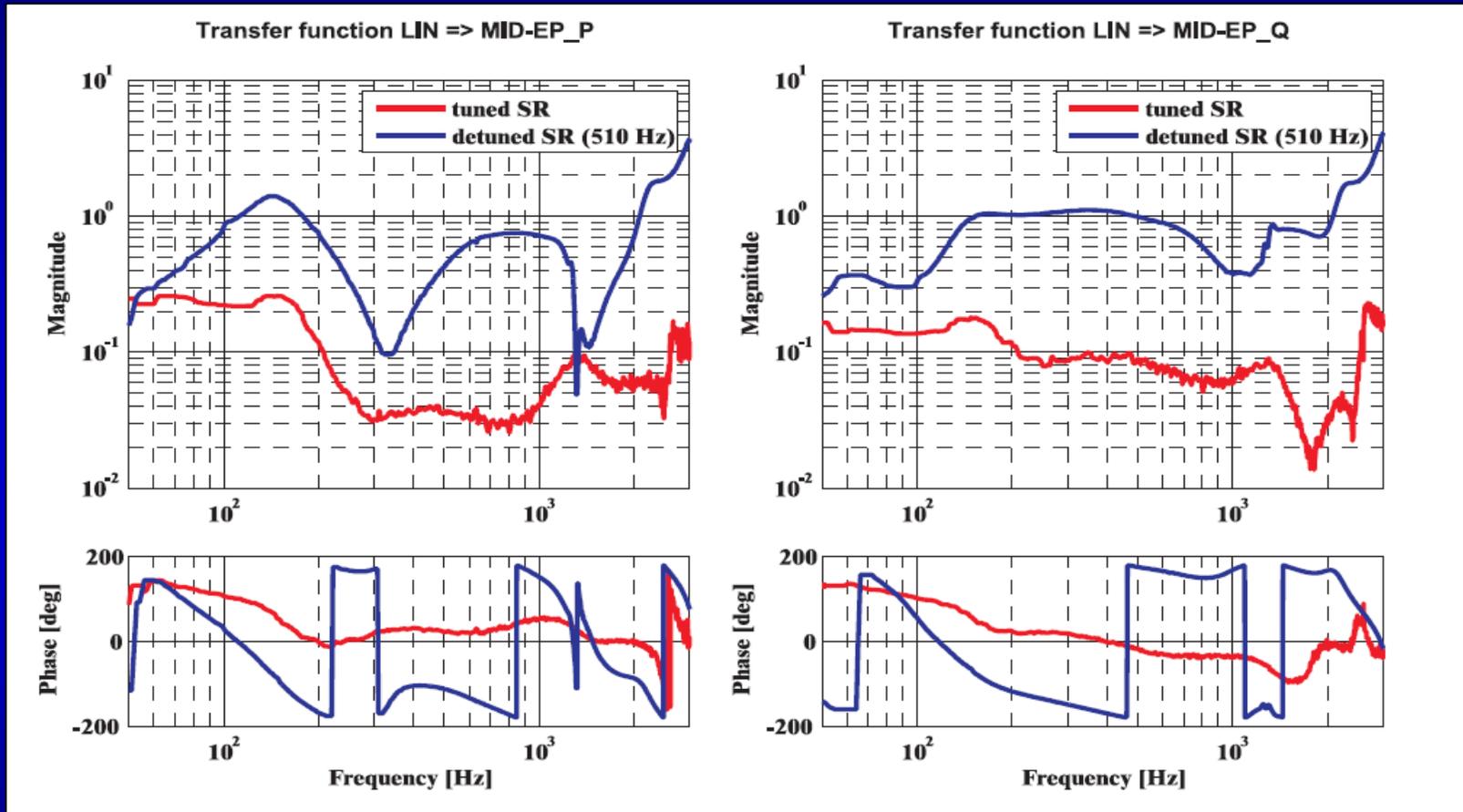
MI Oscillator amplitude noise (OAN)



**OAN coupling @ tuned SR:
P and Q significantly reduced.**



Laser intensity noise (LIN)

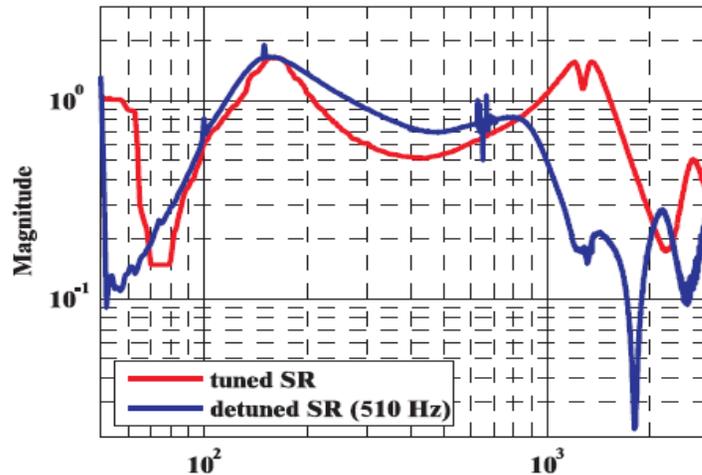


**LIN coupling @ tuned SR:
P and Q significantly reduced.**

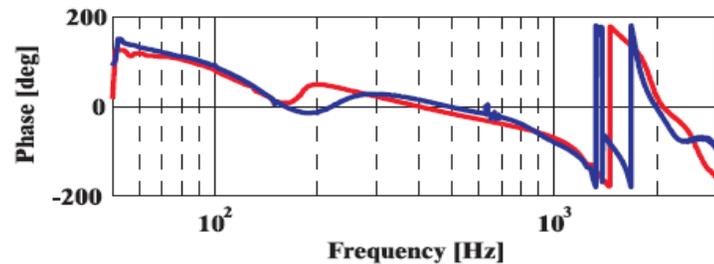
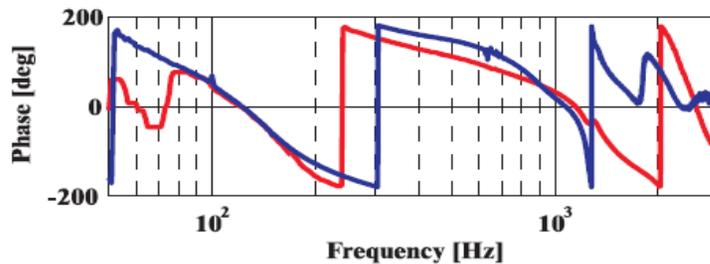
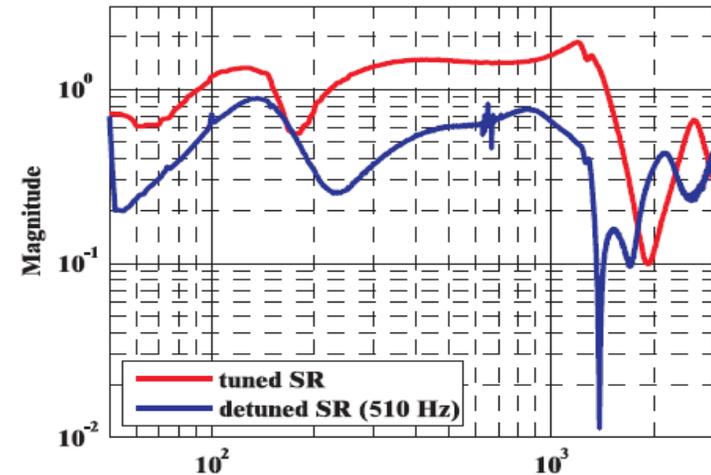


Laser frequency noise

Transfer function Frequency noise => MID-EP_P



Transfer function Frequency noise => MID-EP_Q



**Laser freq noise coupling @ tuned SR:
P and Q worse (but structure slightly simpler).**



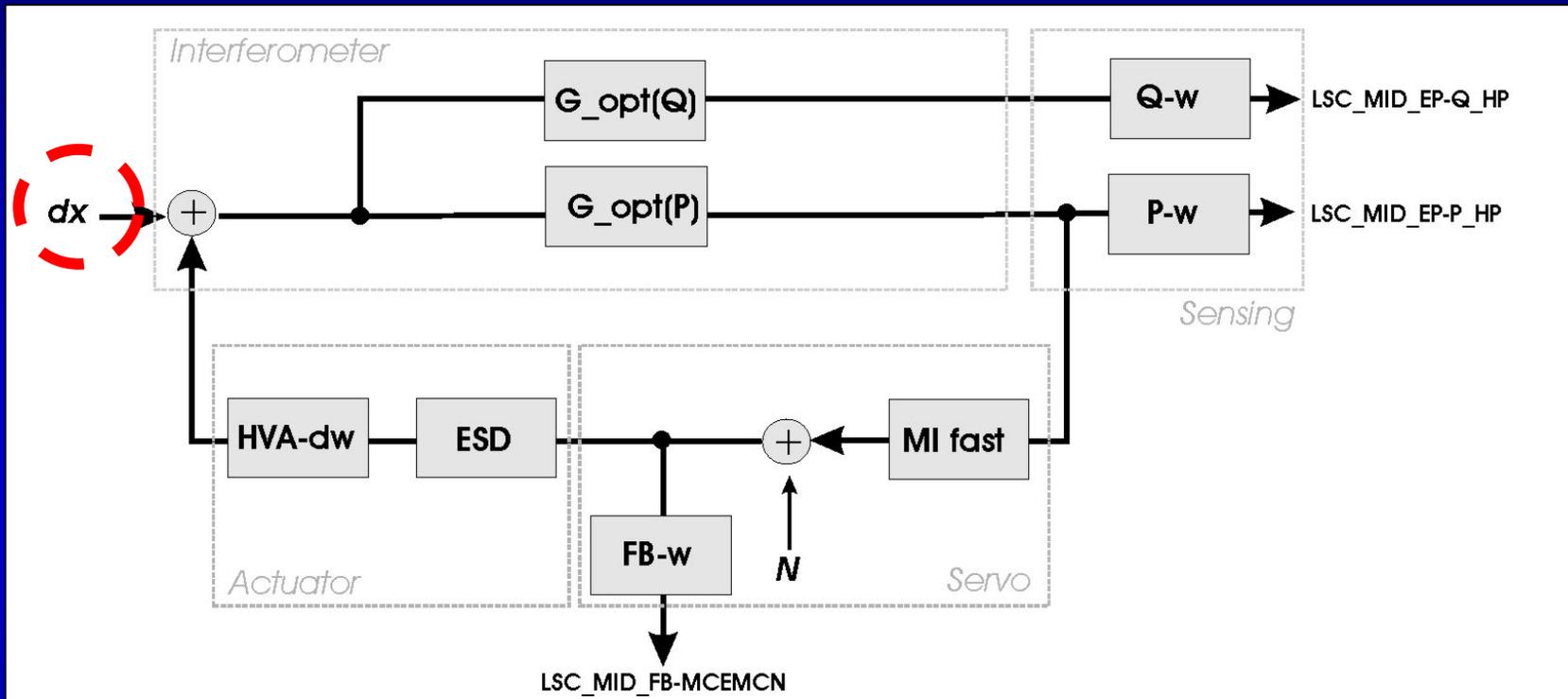
Summary of noise couplings

- OPN coupling @ tuned SR:
P dramatically reduced, Q significantly reduced.
- OAN coupling @ tuned SR:
in both quadratures clearly reduced.
- Laser frequency noise coupling @ tuned SR:
worse in both quadratures.
- Laser intensity noise coupling @ tuned SR:
P and Q significantly reduced.

In general:

The size of TFs for the noise couplings are reduced on average and the structure of the TFs look a bit simpler.

Calibration to starin sensitivity (frequency domain)

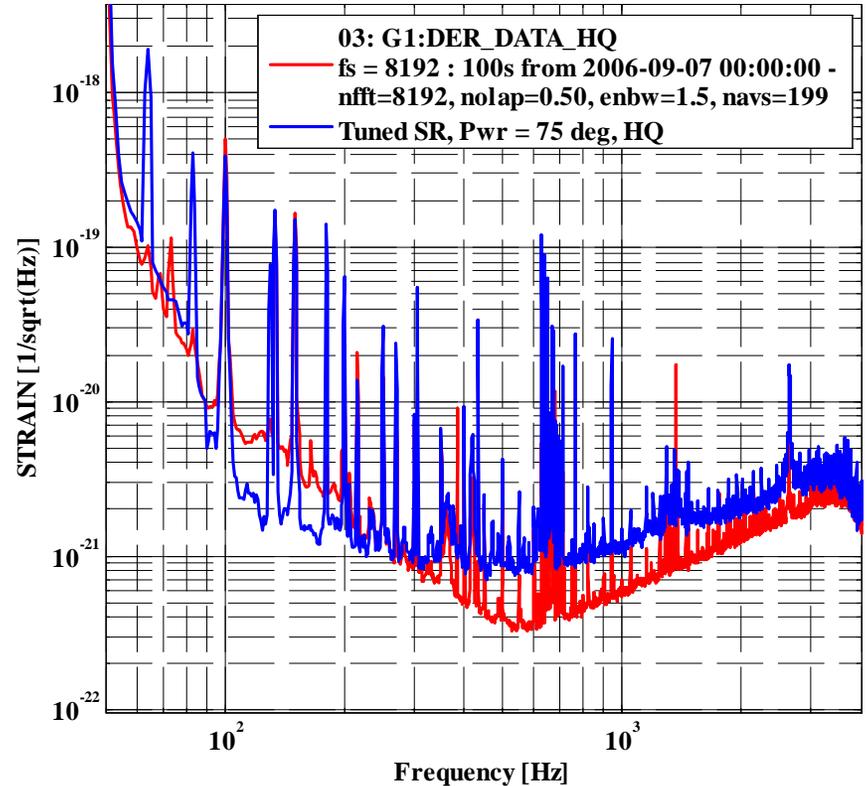
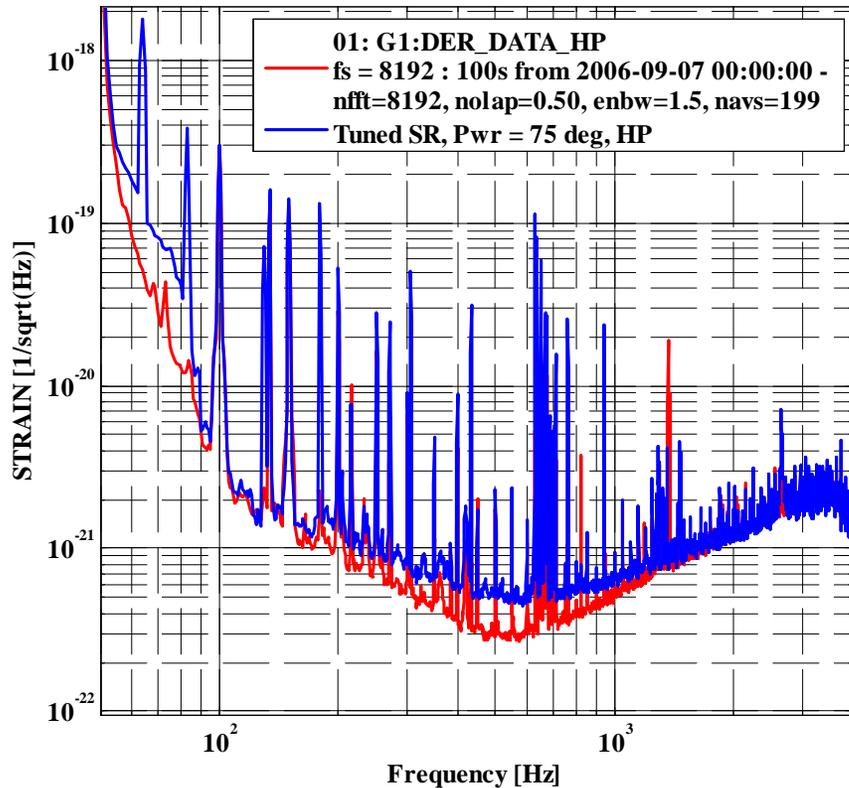


$$dx = \frac{MID_EP-P_HP \cdot P-dw}{G_{opt}(P) \cdot CLTF}$$

$$HP = \frac{dx}{1200\ m'}$$



Sensitivity: tuned vs detuned (550Hz) SR

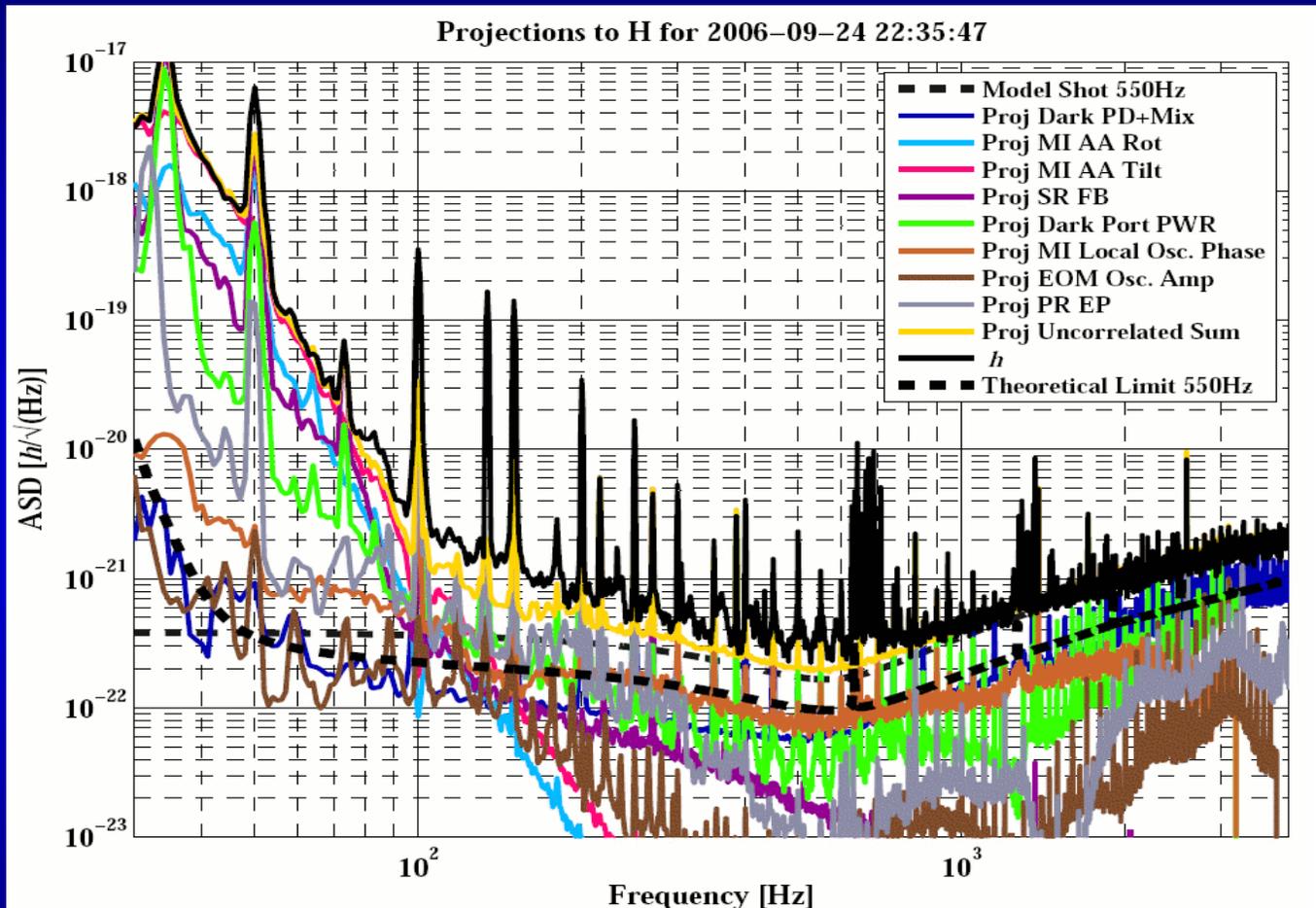


No improvement in sensitivity

(not surprising, as we are not limited by shotnoise at low frequencies and we did not optimize the detector / parameters for tuned SR)



Noise projections for a time of S5



So far we don't understand why Q got worse
at high frequencies for tuned SR. (?)



Summary

- Demonstrated controlled jumping to tuned SR.
- No sensitivity improvement for tuned SR at the moment (spend no time for optimization).
- Some of the noise couplings are significantly reduced for tuned, but their structure is still complicated.
- In tuned case the RF AM is reduced by a factor of 12

Outlook:

To gain understanding of the noise couplings:

- *seems important in order to speed up commissioning of GEO*
- *is essential for good design of next generation of instruments*



E n d

