Scattered Light at GEO 600

Or

``walking right into the cat's eye trap''



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- At the end of September excess noise between 100 Hz and 1 kHz was first noticed.
- All October this noise got worse and was clearly the limiting noise in major part of our detection band.
- At this time we were by far not able to reproduce our best sensitivity from September.

In pricipale we had two candidates:

Scattered light (stray light)



For various reasons we concentrated on scattered light!

Indicators for scattering

- Noise sounds like scattering (listening to MID-EP gives ,wooshes')
- Nonstationarity of noise on short timescale (noise level varies a lot on seconds scale)





Indicators for scattering

- In MID-EPs the noise shape looks like a shoulder
- The noise shoulder is nearly smooth (apart from a few lines no structure or features in it)



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- In S1 we already had problems with stray light (blue curve).
- Now the shape of the MID-EPs is roughly the same.
- In S1 the source of scattering was found to be on the detection bench (and was reduced by replacing optics)



Conclusion: confident that the shoulder was caused by scattering



- Light that leaves the main beam
- Gets a phase shift in respect to main beam
- Reenters main beam and gets detected at the dark port.

<u>Simple picture</u> (without possible (?) enhancement in SR-cavity): To get scattered light at 1 kHz we need a phase shift of roughly $2000\pi/s$.

Two possiblities:

- Component moving at low frequency about many fringes = 1mm/s
- Component moving a little but very fast (1kHz)

We worry just about a very small light level !

PROBLEM: NONLINEARITY

Scattering in common and diffrential path



Scattering in common path path would be less ``effective'' than in the differential path !

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Identifying the scattering source

The filter experiment:

- Putting an absorbtion Filter (T=50%) in the path between IFO and potential scattering source
- Light amplitude at source and scattering will be reduced by a factor sqrt(2)
- Scattered light amplitude going back into IFO is attenuated by sqrt(2) again at filter
- In summary an exact factor of 2 less scattered light in IFO





Possible differential scattering points

B → Transmission through the inboard mirrors: scattering from ESDs?



No problem so far !

 $C \rightarrow$ Both beams seem to be properly dumped !



Scattering from the endstations

By listening to MID-EPs we could already hear a noticeable influence from tapping IR-Filter and CCD.

But wasn't the origin of the shoulder !











Used a filter (located near the waist) to intentionally force



bd2

bd1

Changed the optical setup at SR-bench



- Tried different alignments through L1 and L2
- Replaced bd2 with ionsputtered mirror

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- Replaced the scanner (poor quality mirrors) with a mirror and removed second lens

Changed the optical setup at SR-bench



- Tried different alignments through L1 and L2
- Replaced bd2 with ionsputtered mirror
- Replaced the scanner (poor quality mirrors) with a mirror and removed second lens
- Ended up with a very simple setup:
 - only two components left
 - no waist !

Known amplitude noise coupling

We are still not sure about the coupling path 🛞 !!

Unfortunately our detector knowledge from this time was limited by a DAQs crash and consequences:

- didn't record important channels
- no noise projections for nearly three months

One exemplary explanation of the shoulder: amplitude noise coupling







Due to the changes at the signal recycling bench we were able to improve sensitivity by a factor of 3 to 4 from 100 Hz to 1 kHz.

Since then we had no indication for any further scattering \Rightarrow could carry on commissioning and improving sensitivity as before !!

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- Try to avoid beam waist wherever it is possible
- Especially in differential beam paths

If you can't avoid having a waist:

- Don't place any optics near the waist !
- Use high quality optics

And be always aware of the cat's eye trap!

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